

*Cosmological Insights from  
Gravitational Lensing – or: how Herr  
Doktor Einstein Can Come to the  
Rescue and Help Us Solve the  
Grandest Mystery in  
Modern-Day Physics*

Talk to the Night Sky Network regular meeting  
(arranged by Dr. David Prosper)

Th Sept 25 2014

Dr. Mandeep S.S. Gill

Observational Cosmologist at KIPAC,  
SLAC National Accelerator Laboratory (SLAC NAL), Stanford, CA.

For more info, see also “Extra Cosmo Resources” link near top of:

**Mandeep.Org**

And if you like the talk feel free to “Like” it and/or ask more questions directly at:

<https://www.facebook.com/MSSGTalks>

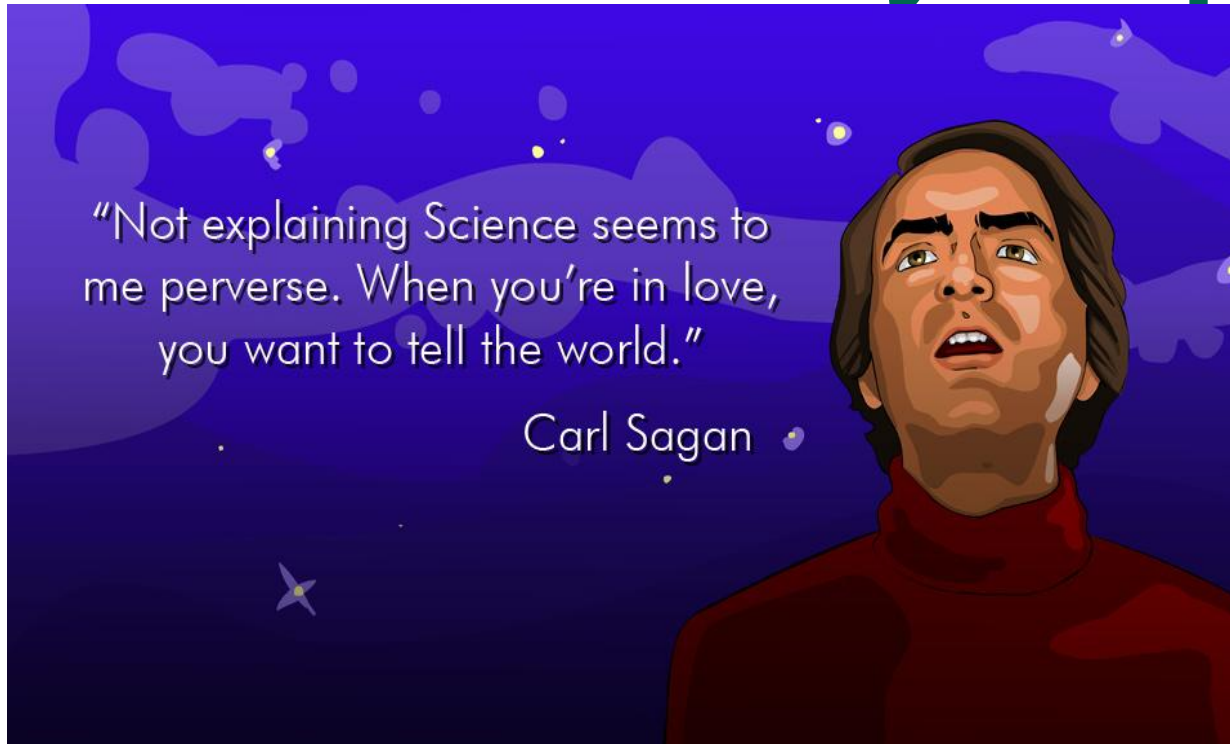
For more information/further questions : [me@mandeep.org](mailto:me@mandeep.org)

(Also listed at the bottom of ***mandeep.org***)

# Talk Plan

- The connections of the smallest with the largest
- “Dark Matter” - DM
- “Dark Energy” - DE
- Bring all this back to Earth for a minute
- Introducing gravitational lensing
- How lensing can help us with DM and DE determination
- The future is bright in this field!

# Some of My Inspirations

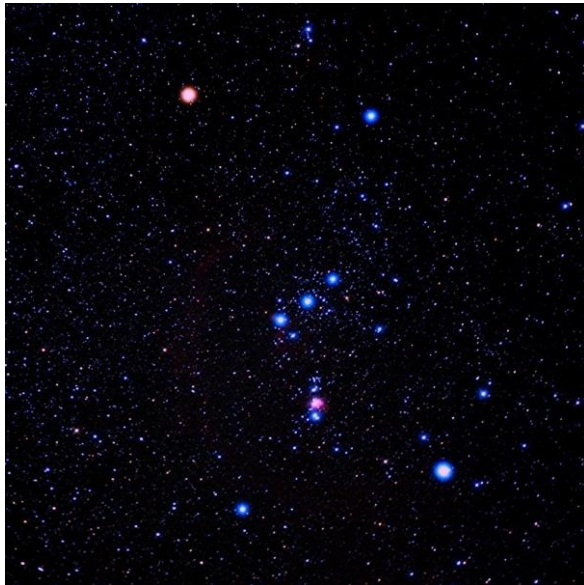


"Not explaining Science seems to me perverse. When you're in love, you want to tell the world."

Carl Sagan

Great awe-filled  
scientists of the past

Looking up  
at the Night  
Sky



Being in  
Nature





# One thing to take away



After evolving separately for decades, particle physics and cosmology have strongly reconverged

Essentially and what the Universe is made of, and how it came to be, are inextricably linked ideas at root.



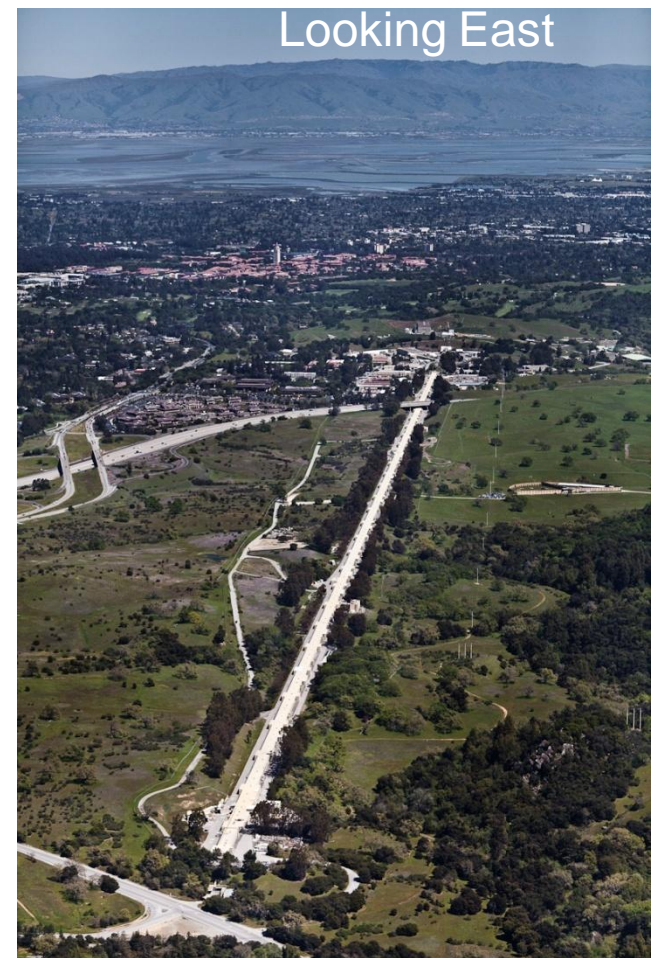


# SLAC NAL

Looking West



Looking East



Campus still looks just fine





# Astrophysics

Studying the very large requires scaling up your tools in a very large way.

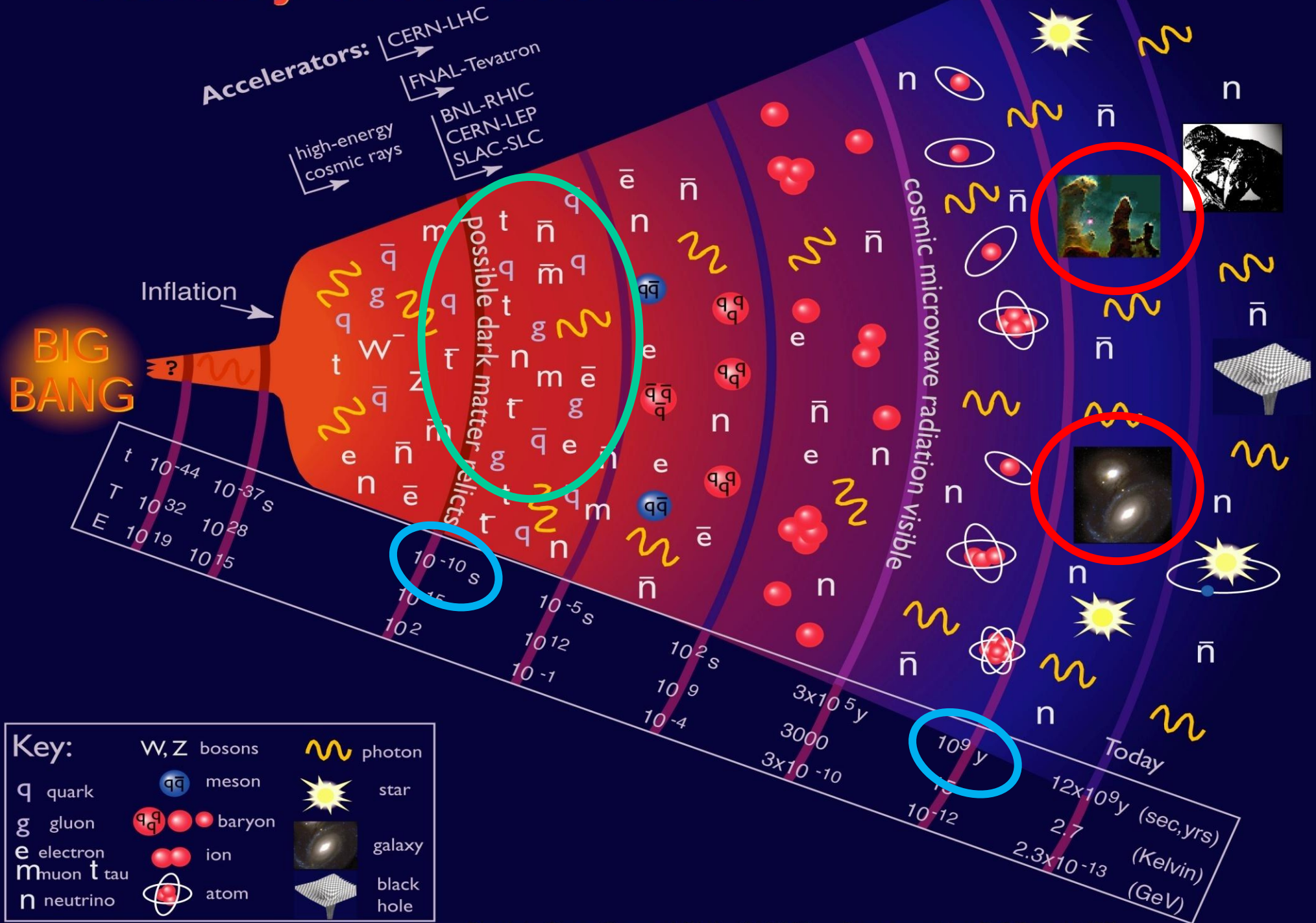
FERMI: Gamma-ray



Chandra: X-ray



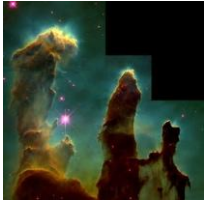
# History of the Universe in one slide!





But – much more than that to our current picture of the Universe, which is actually kind of... *Cupcake-like!*

Sprinkles =  
Normal  
Matter: ~4%



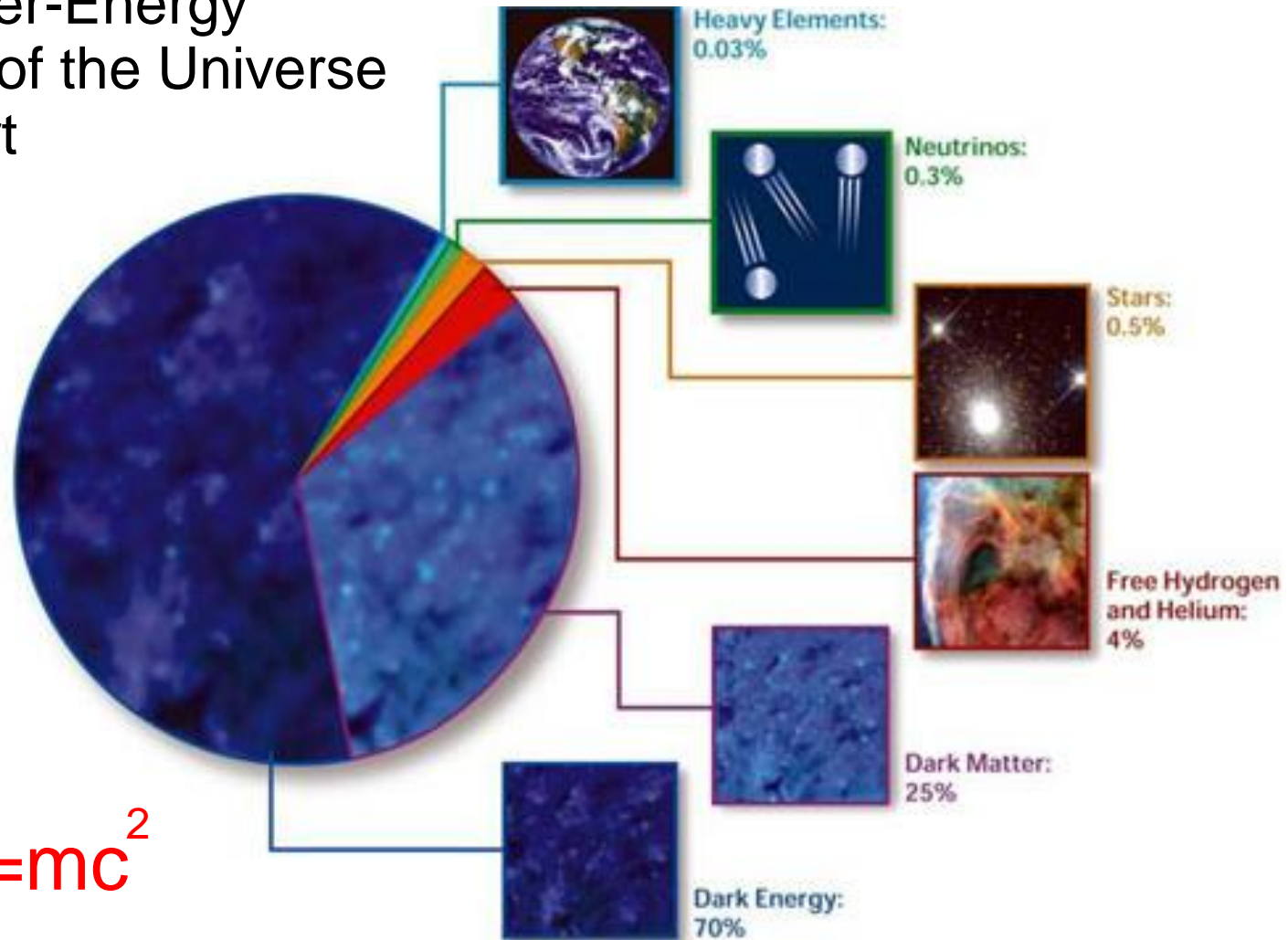
Frosting =  
Dark matter: ~26%



Main Cupcake:  
Dark energy: ~70%

# The biggest mystery in all science..??!

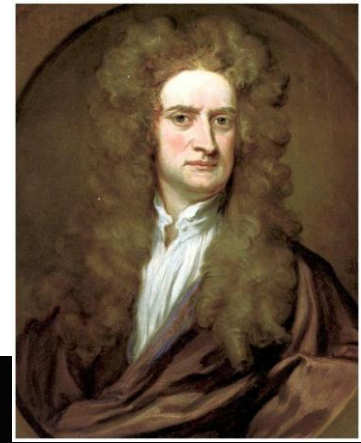
Shown: Matter-Energy  
composition of the Universe  
as a pie chart



Implicit here:  $E=mc^2$



# Why do we believe in DM? Let's start in 1689: Newton explains the motion of the planets



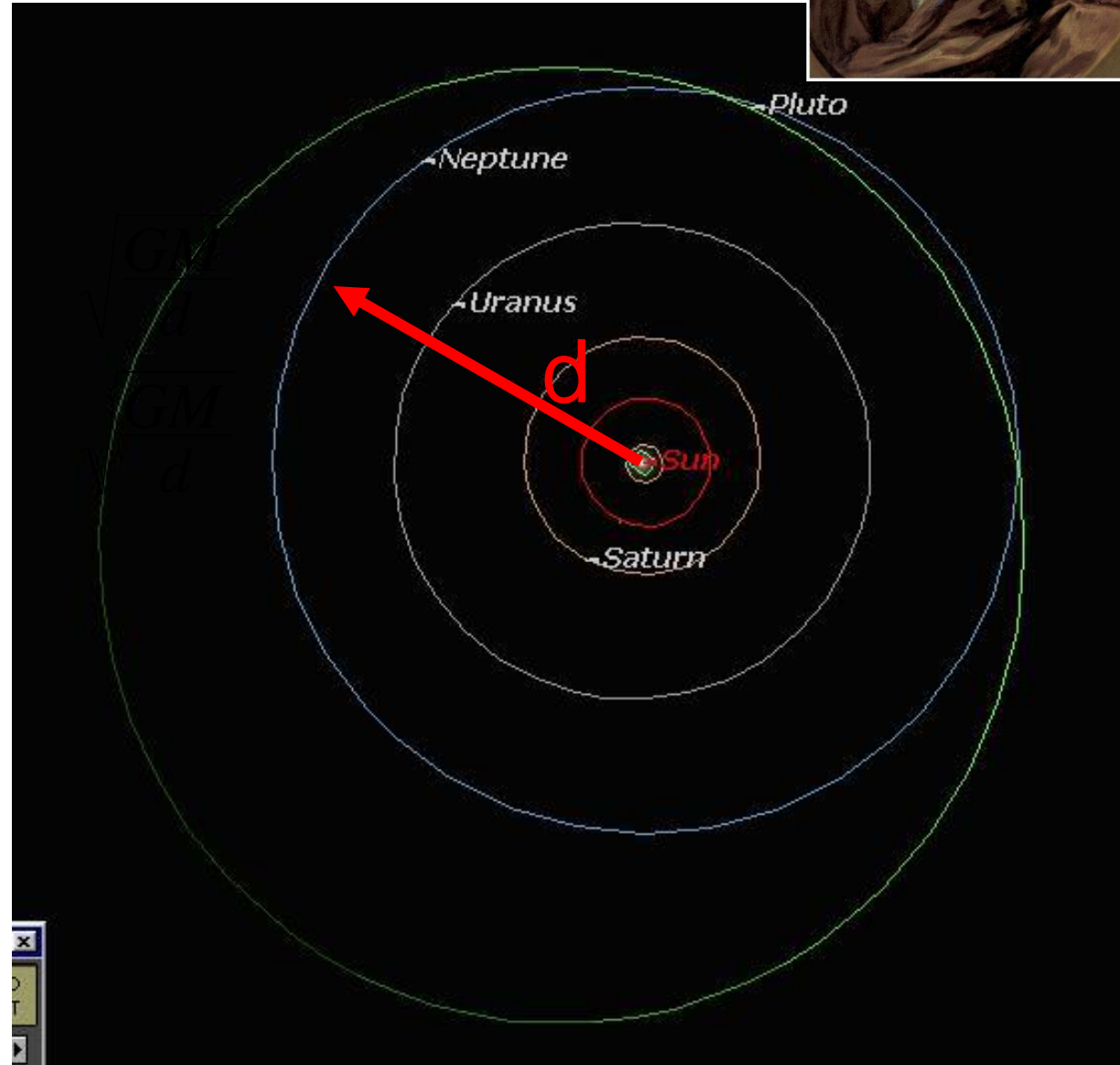
Matter produces force  
(Universal Gravity)

Force causes acceleration

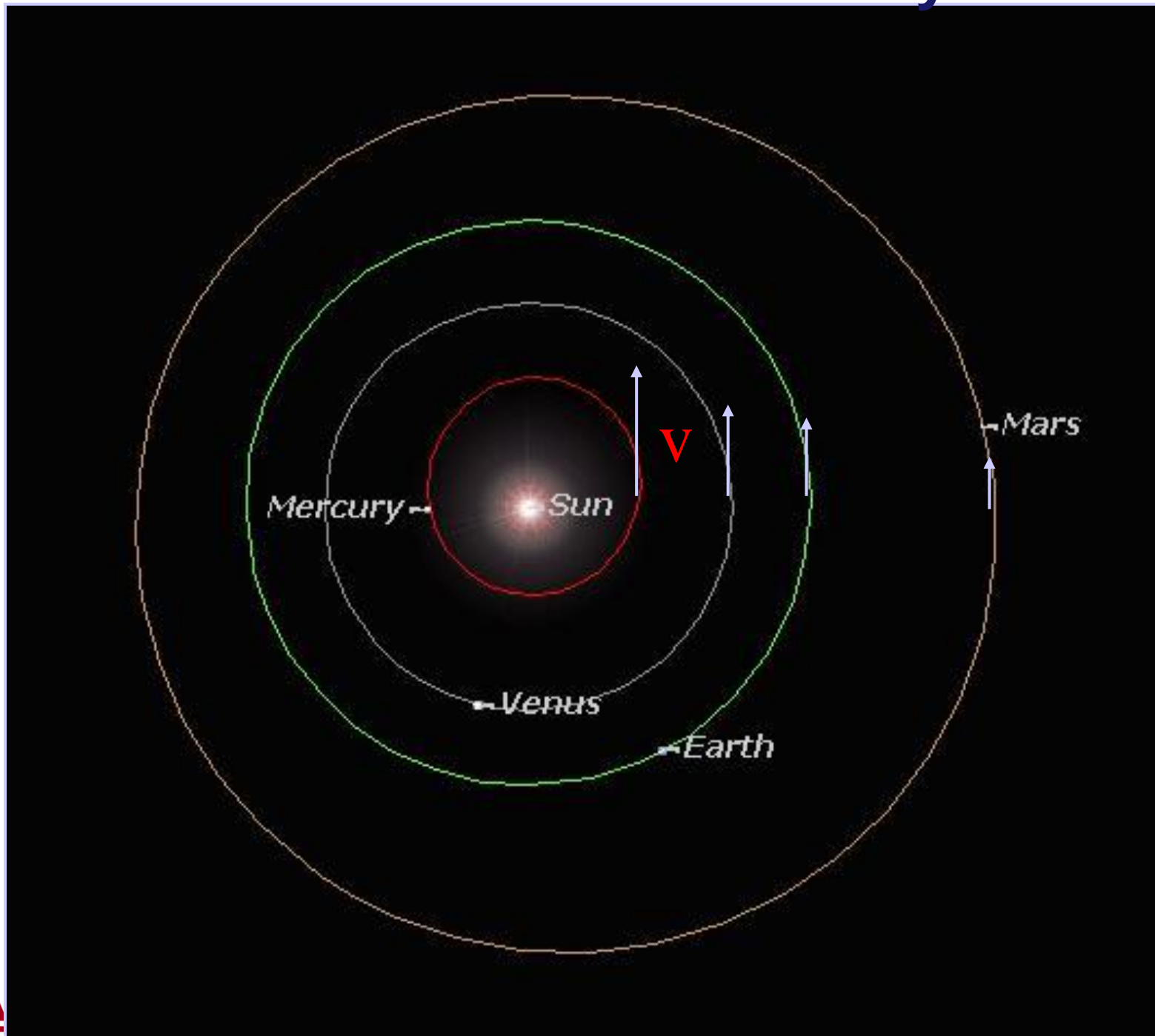
$$F = \frac{GMm}{d^2}$$

$$a = \frac{F}{m} = \frac{v^2}{d}$$

$$v = \sqrt{\frac{GM}{d}}$$



# And we see in the Solar System:

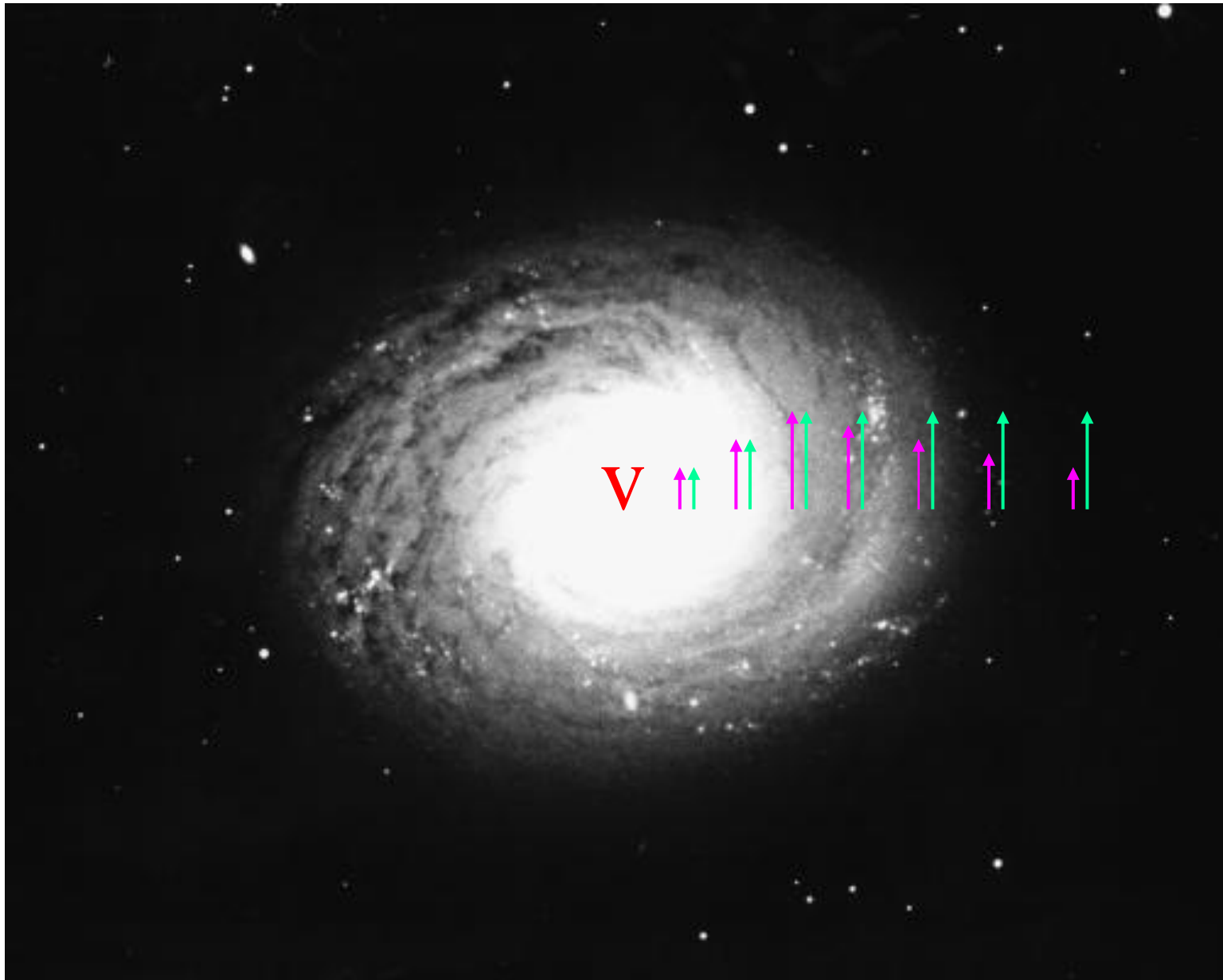




# In Galaxies we expected:



# But we measure (green arrows):



!!!!



# DM and Galaxy Stellar Orbital Speeds

- Thus we infer there's some other kind of matter we can't see in visible light *directly*.
- Evidence for it is all from gravity:
  - Velocity (measured by Doppler Shift) of stars and gas clouds imply mass of galaxy continues beyond visible edge.

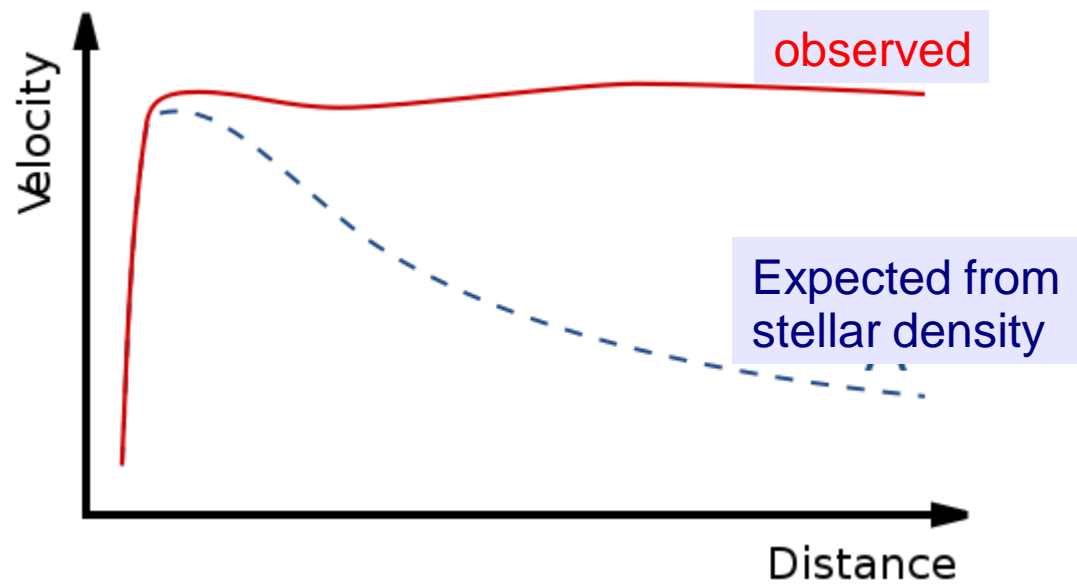
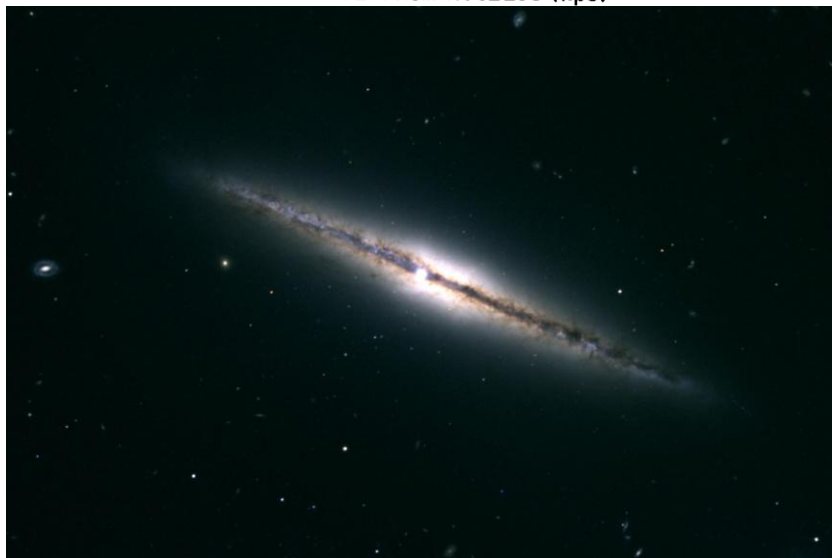
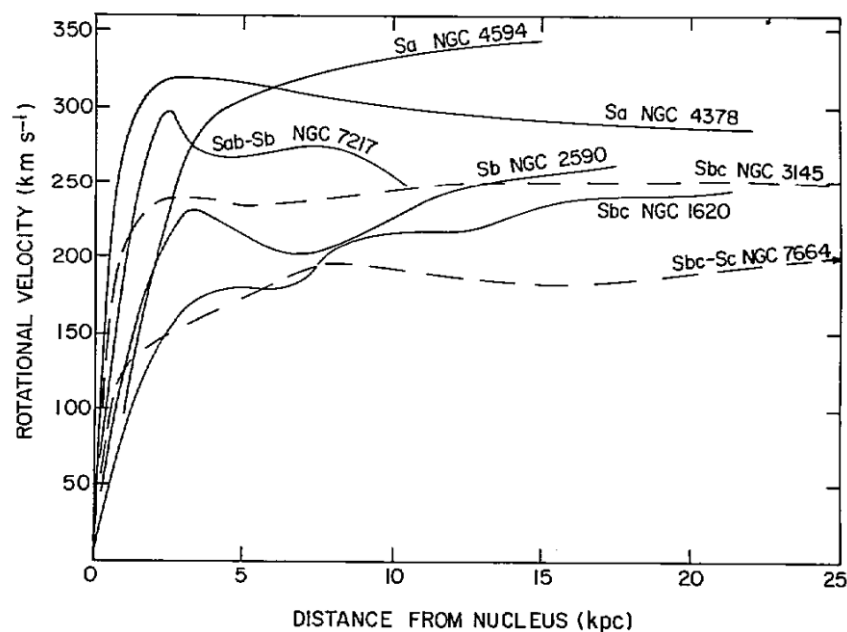


Image credit: wikipedia

Distance from center of galaxy

Mid-1970s -- Rubin, Ford, Roberts, and others saw that:  
Spiral galaxies have flat rotation curves



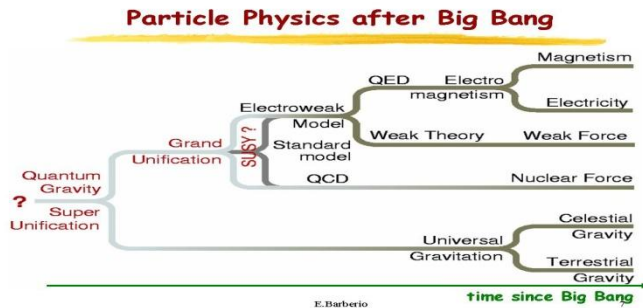
# Dark Matter is indirectly “seen” in many arenas

## Galaxy Dynamics &:

- Galaxy cluster dynamics
- Gravitational Lensing
- Large Scale Structure
- Baryon Acoustic Oscillations
- Cosmic Microwave Background



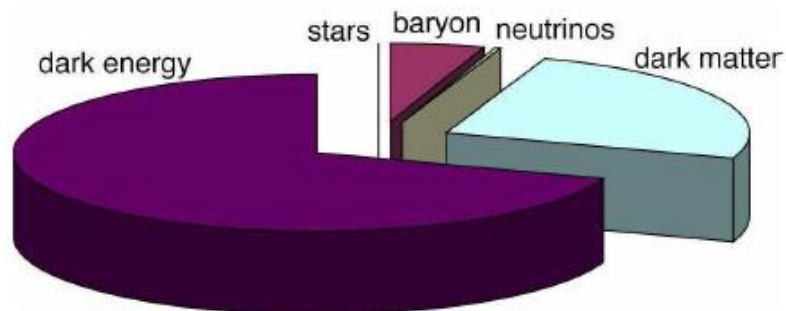
## Particle Theory



And searched for in others: Colliders  
+ direct DM detection expts



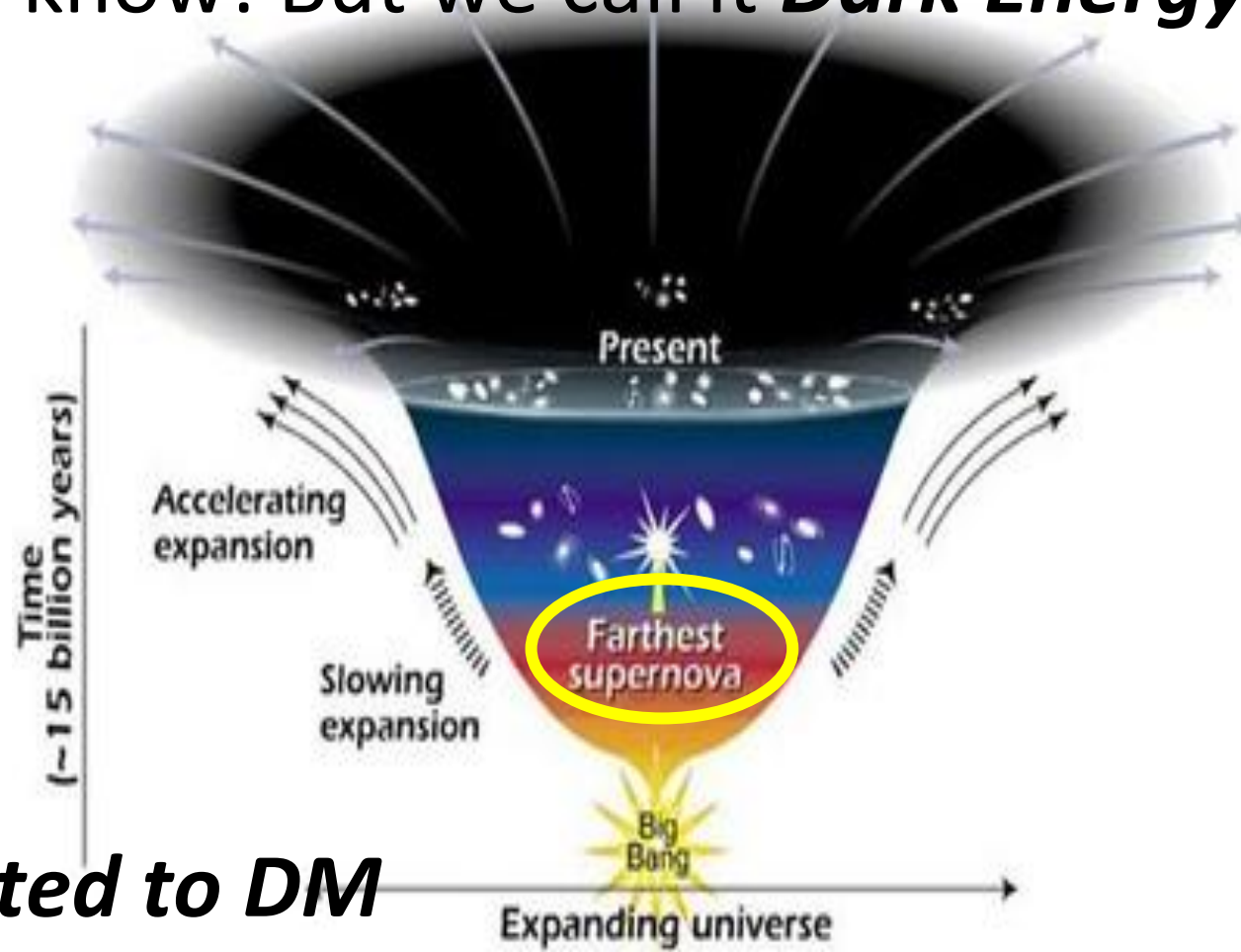
## Cosmological Component





# Next: What's Dark Energy (DE) ?

- 1998: Discovery that expansion of the universe is accelerating... what's driving this? (c.f. 2012 Nobel Prize in Physics)
- We don't know! But we call it ***Dark Energy***.



- ***NOT*** related to DM

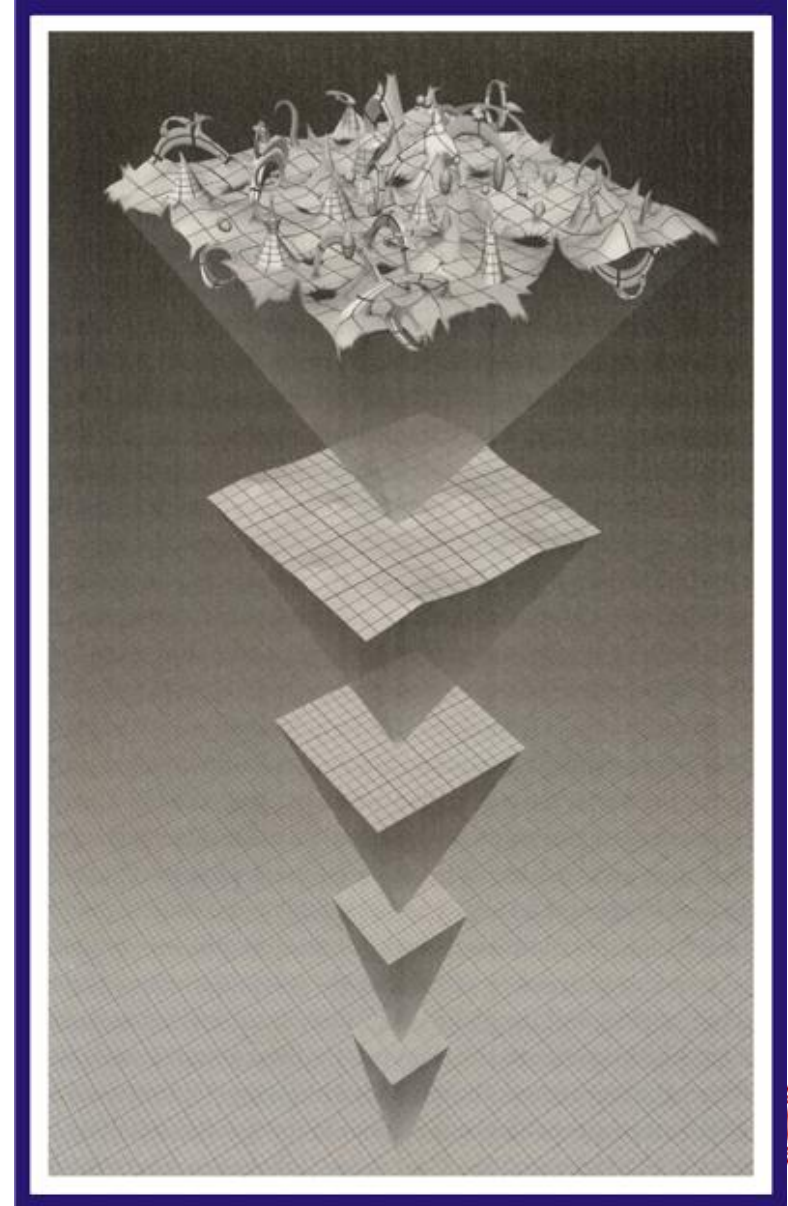
# At smallest scales, vacuum is a writhing mass of particles

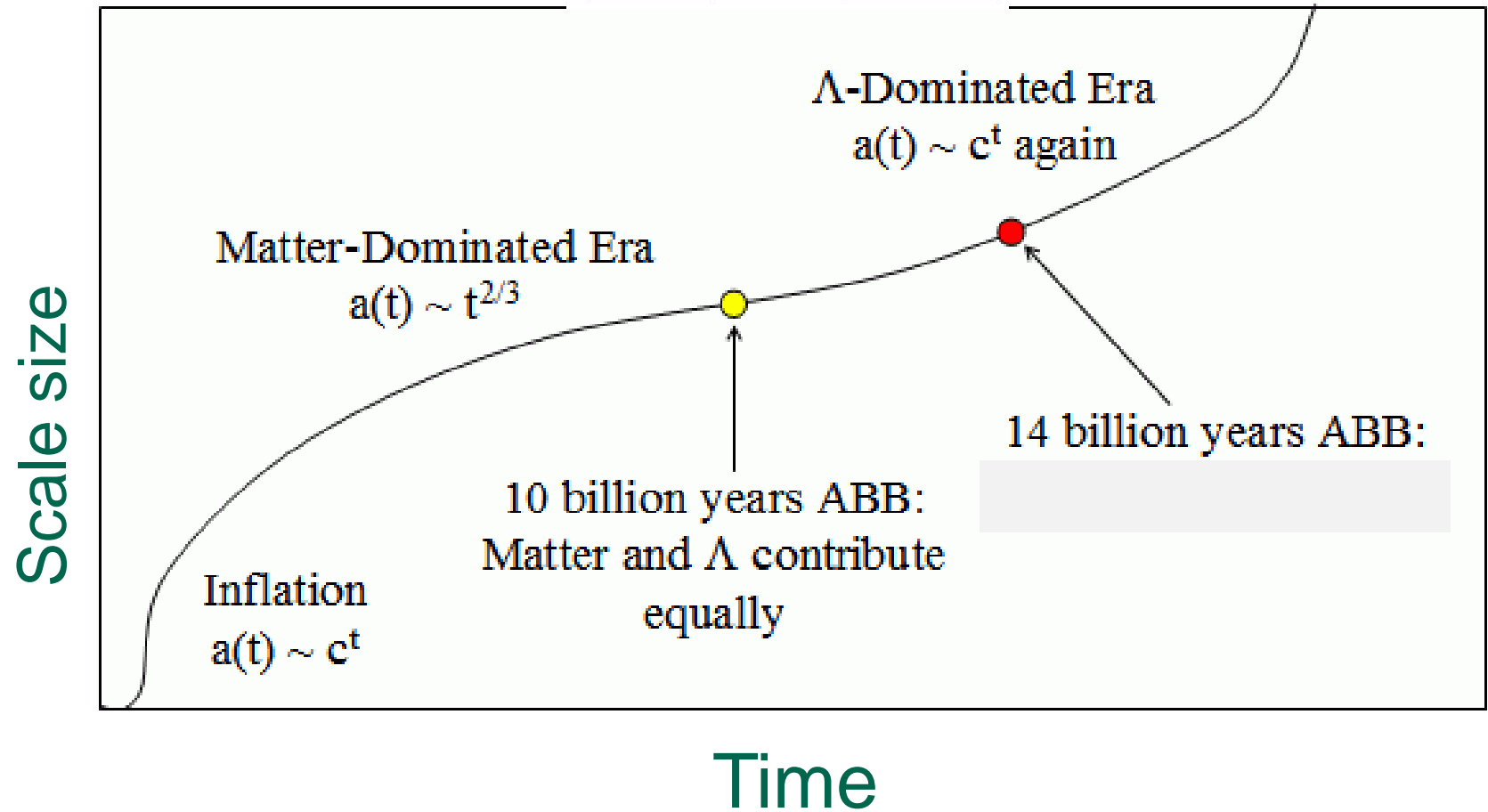
We already see this at some level (cf. *Casimir Effect*) for charged particles.

But if spacetime becomes “foamy” at the Planck Length, then we expect the vacuum to have  $10^{128}$  (!!!) times as much energy as is seen.

[For more info, see also “Extra Cosmo Resources” link near top of:

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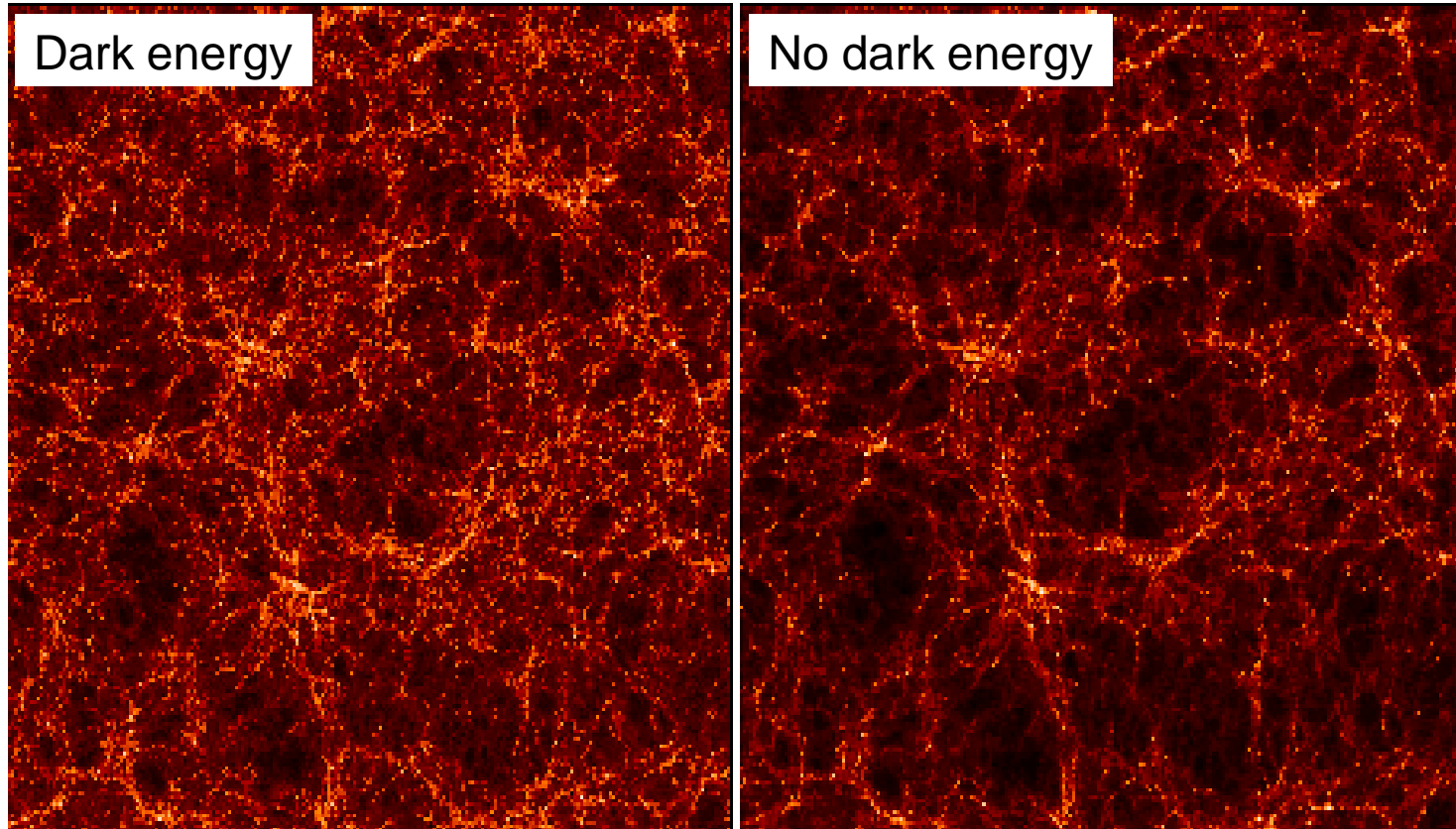






# Another place we see DE: Simulations of Large Scale Structure of matter in the universe

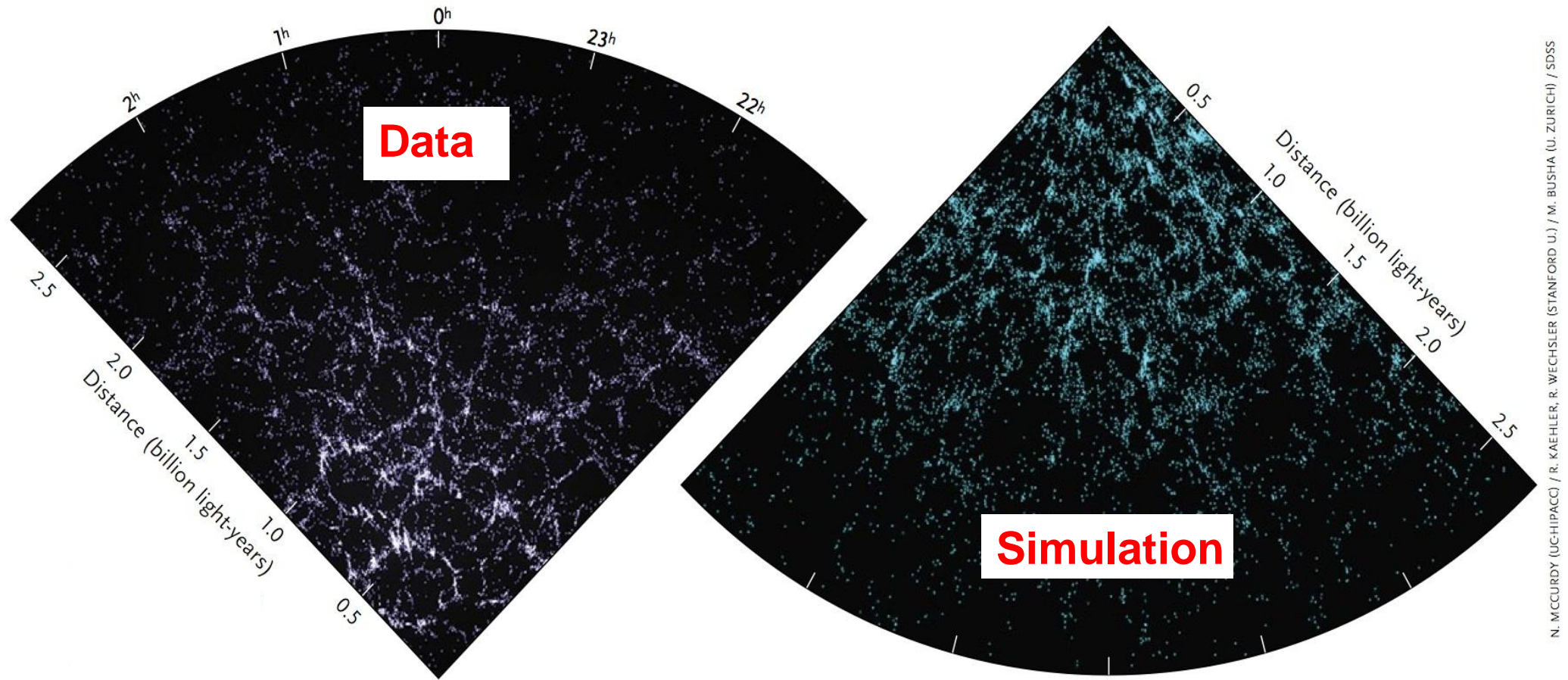
## 12 billion years ago



The VIRGO Collaboration 1996

[Show simIn]

# Now compare Data vs. Simulation – very similar

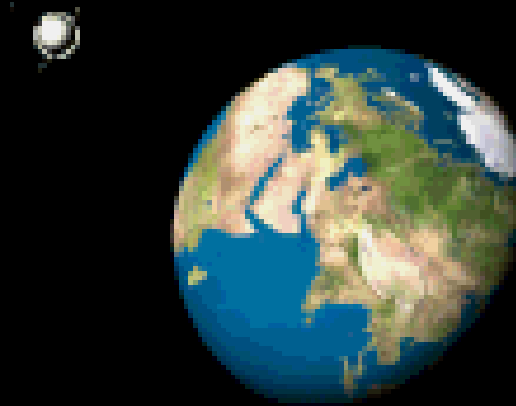


Simulation of a universe with 23% dark matter and 72% dark energy

# How are we ever going to figure this out..?

First let's step back to what we know. For a sense of scale:  
begin close to home with travels around the Earth's  
circumference

8,000 km



Fastest Low Earth Orbit Satellites: ~30,000 kph

Elapsed Time for Orbit: ~90 min

How about for Light?

**1/8 of a second!**



# From 'Near' to Far:

## Step out in distances:

Earth's circ. – 0.13 light sec

the Moon -- 1.3 light sec

Sun -- 8.3 light min

Mars -- 14 light min

Neptune -- 5 light hrs

Pluto -- 6 light hrs

Voyager 1 -- 17 light hrs

Oort Cloud - ~5 light months

Alpha Centauri System - ~4 lyr

Solar Nbhd - ~20 lyr

Big Dipper Stars - ~70-120 lyr

Betelgeuse and Rigel - ~650, 860 lyr

MWG center - ~30 klyr

Andromeda - ~2.6 Mlyr

Nearby galaxies - ~10-100 Mlyr

Farthest visible objects - ~13 Glyr

→ Limit of Observable Universe:

CMB (Cosmic Microwave Background) - 13.8 Glyr

# Gravitational Lensing

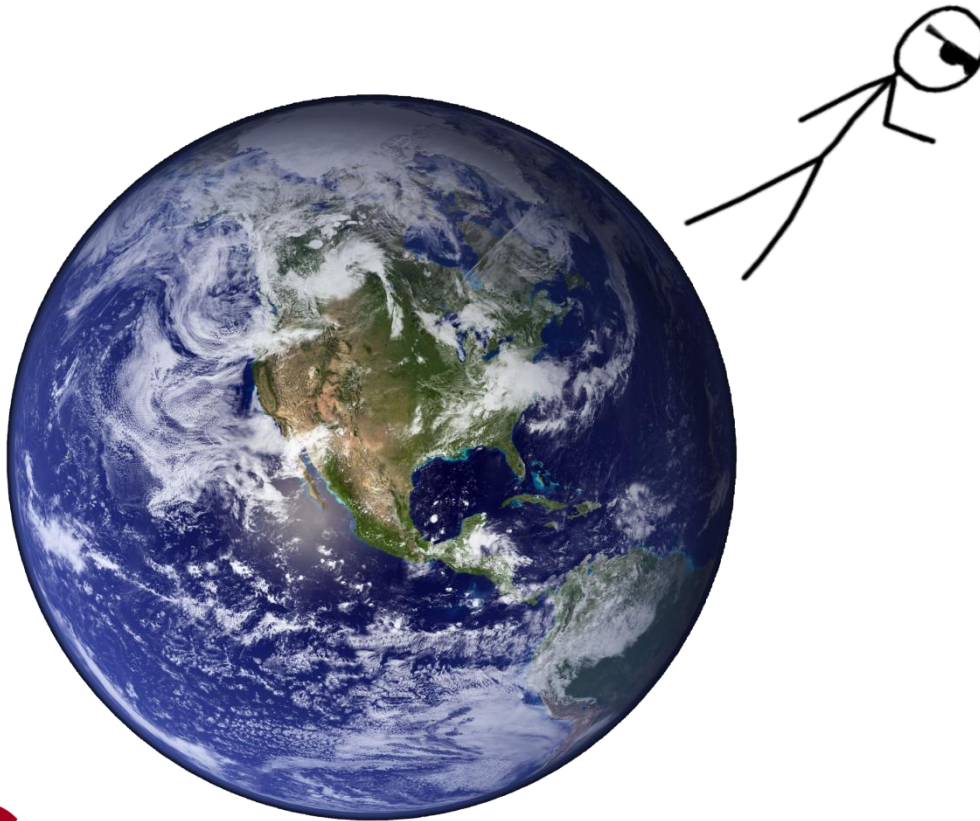
The background of the slide is a high-resolution astronomical image, likely from the Hubble Space Telescope, showing a dense field of galaxies. The galaxies appear as bright, yellowish-white points of light, some with visible spiral or elliptical structures, set against a dark, almost black, cosmic background. The distribution of galaxies is uneven, with some areas being more crowded than others, illustrating the large-scale structure of the universe.

- History of the idea
- What is it?
- Lensing and Dark Matter
- Lensing and Dark Energy

# What does gravity do?

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- Holds us on to the Earth...





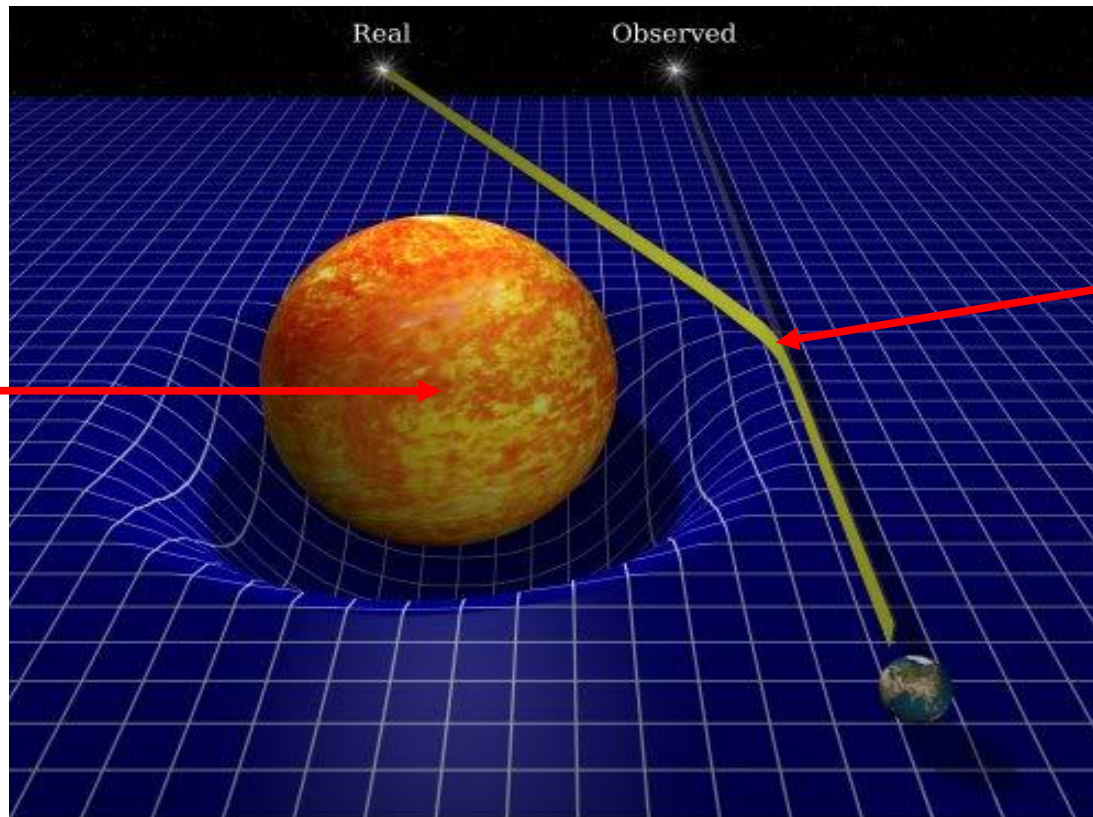
# What does gravity do?

---

- Holds us on to the Earth...
- Holds the Earth in orbit around the Sun...

# What does gravity do?

- Holds us on to the Earth...
- Holds the Earth in orbit around the Sun...
- And *bends light!*



Massive object →  
large gravitational  
field

Light ray is  
deflected by  
gravitational  
field

# The History of an Idea

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- 1687: Newton's Law of Gravity



**Gravity.**  
It's not just a good idea.  
It's the Law.

- 1804: if something is heavy enough, can it deflect light itself? (J. von Soldner)



# The History of an Idea

---

- 1911: Einstein calculated effect of gravity on light rays

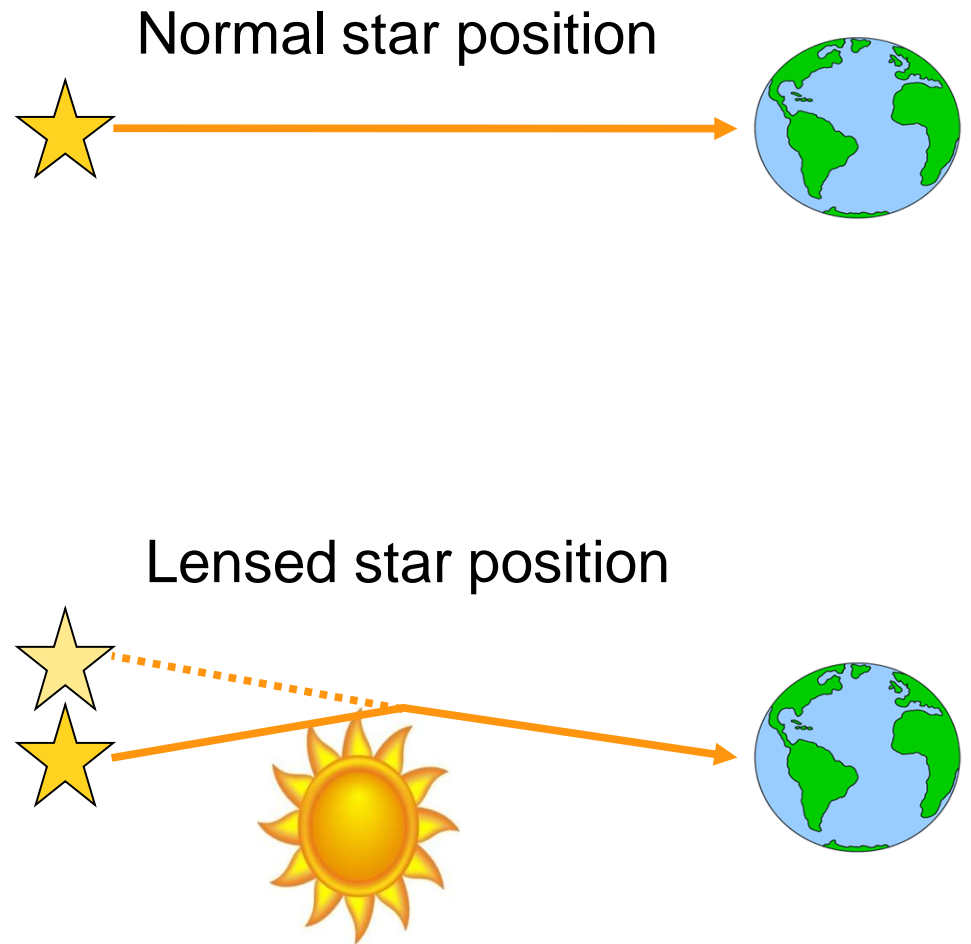
(Factor of 2 relative to Newton)



# The History of an Idea

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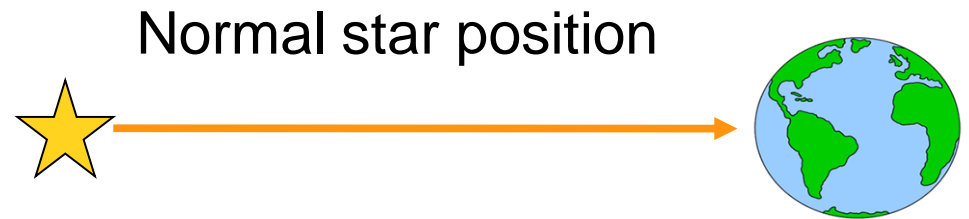
1911: Einstein calculated  
effect of gravity on light  
rays



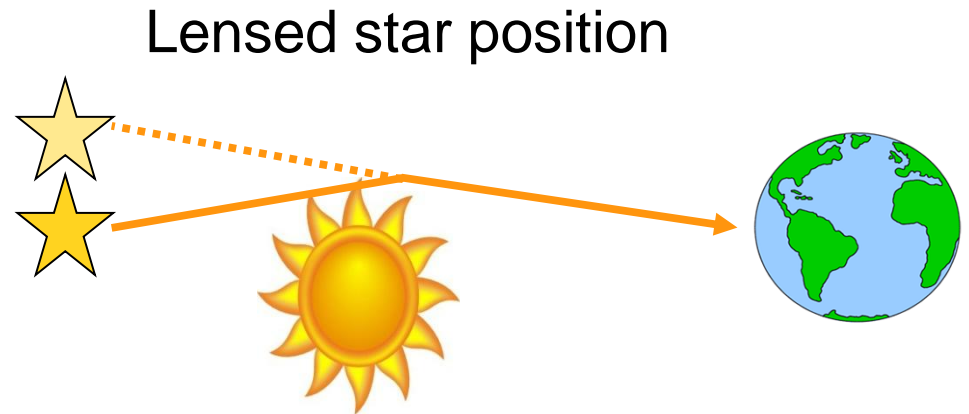
# The History of an Idea

---

1911: Einstein calculated effect of gravity on light rays



- 1919: light observed from star deflected by the Sun during eclipse



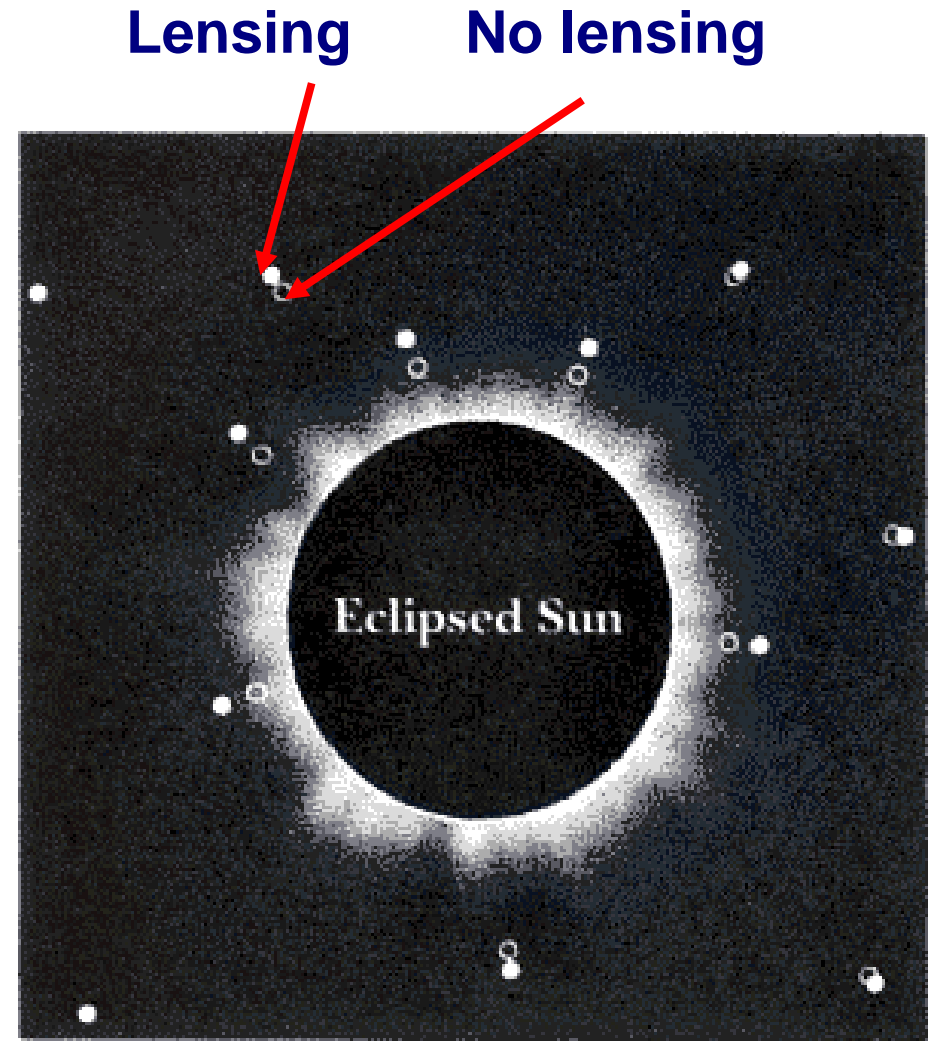


# The History of an Idea

---

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# The History of an Idea

---

- Development of idea that gravity can act as a 'lens'
- In 1936 Einstein described the more dramatic lensing effects (multiple images/ring) and concluded...

# The History of an Idea

---

- Development of idea that gravity can act as a 'lens'
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## DISCUSSION

### LENS-LIKE ACTION OF A STAR BY THE DEVIATION OF LIGHT IN THE GRAVITATIONAL FIELD

SOME time ago, R. W. Mandl paid me a visit and asked me to publish the results of a little calculation which I had made at his request. This note complies

not decrease like  $1/D$ , but like  $1/\sqrt{D}$ , as the distance  $D$  increases

Of course, there is no hope of observing this phenomenon directly. First, we shall scarcely ever approach closely enough to such a central line. Second, the angle  $\beta$  will defy the resolving power of our



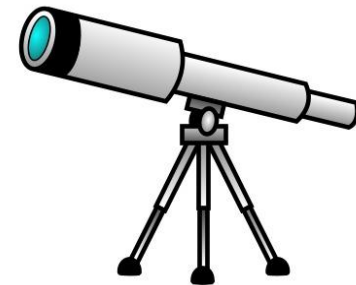
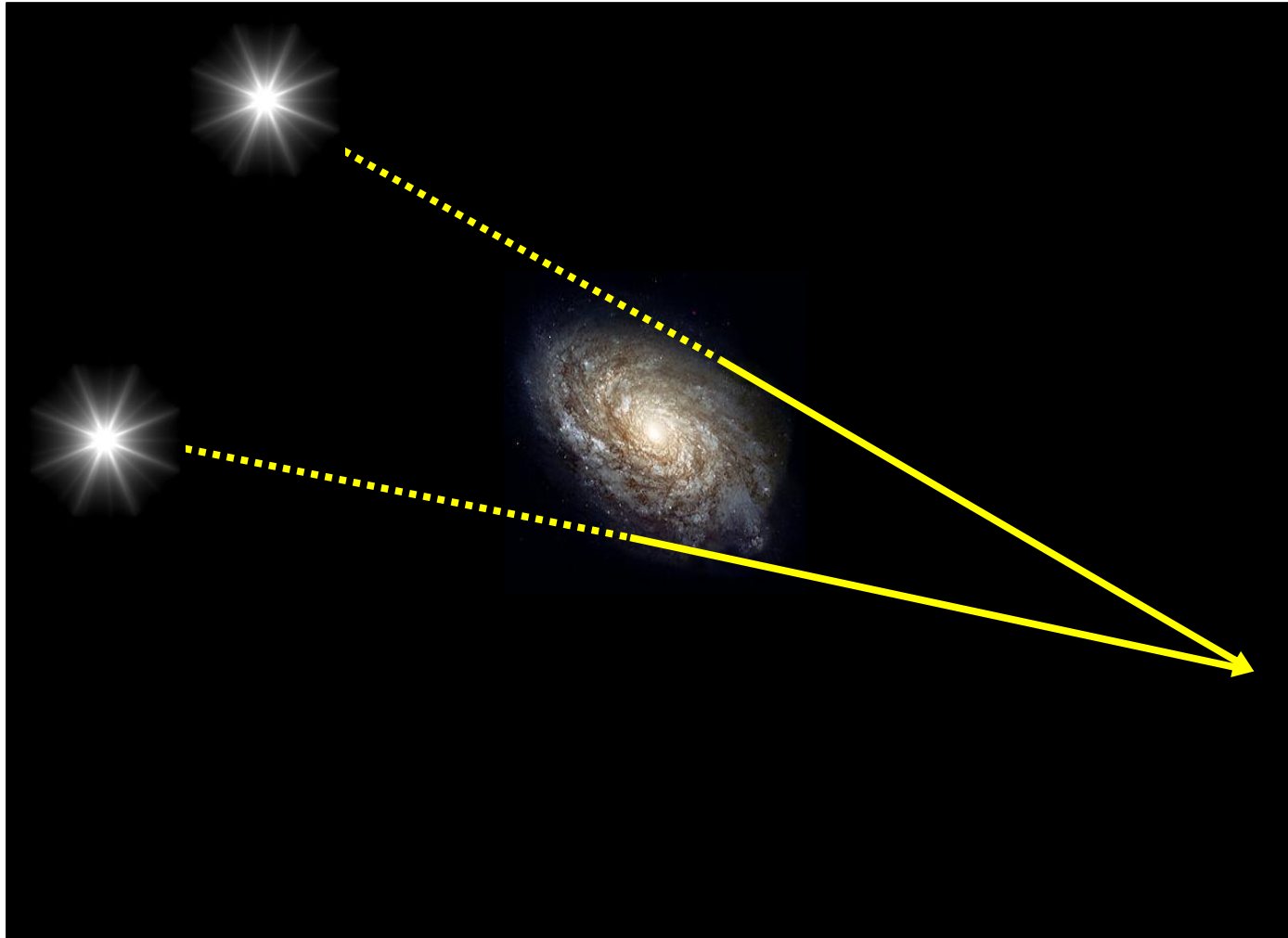
# Gravity *lenses* light with several effects

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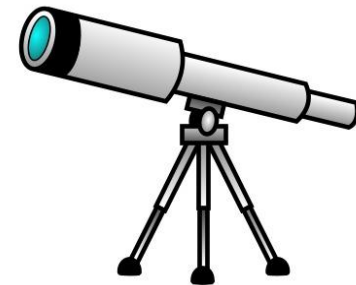
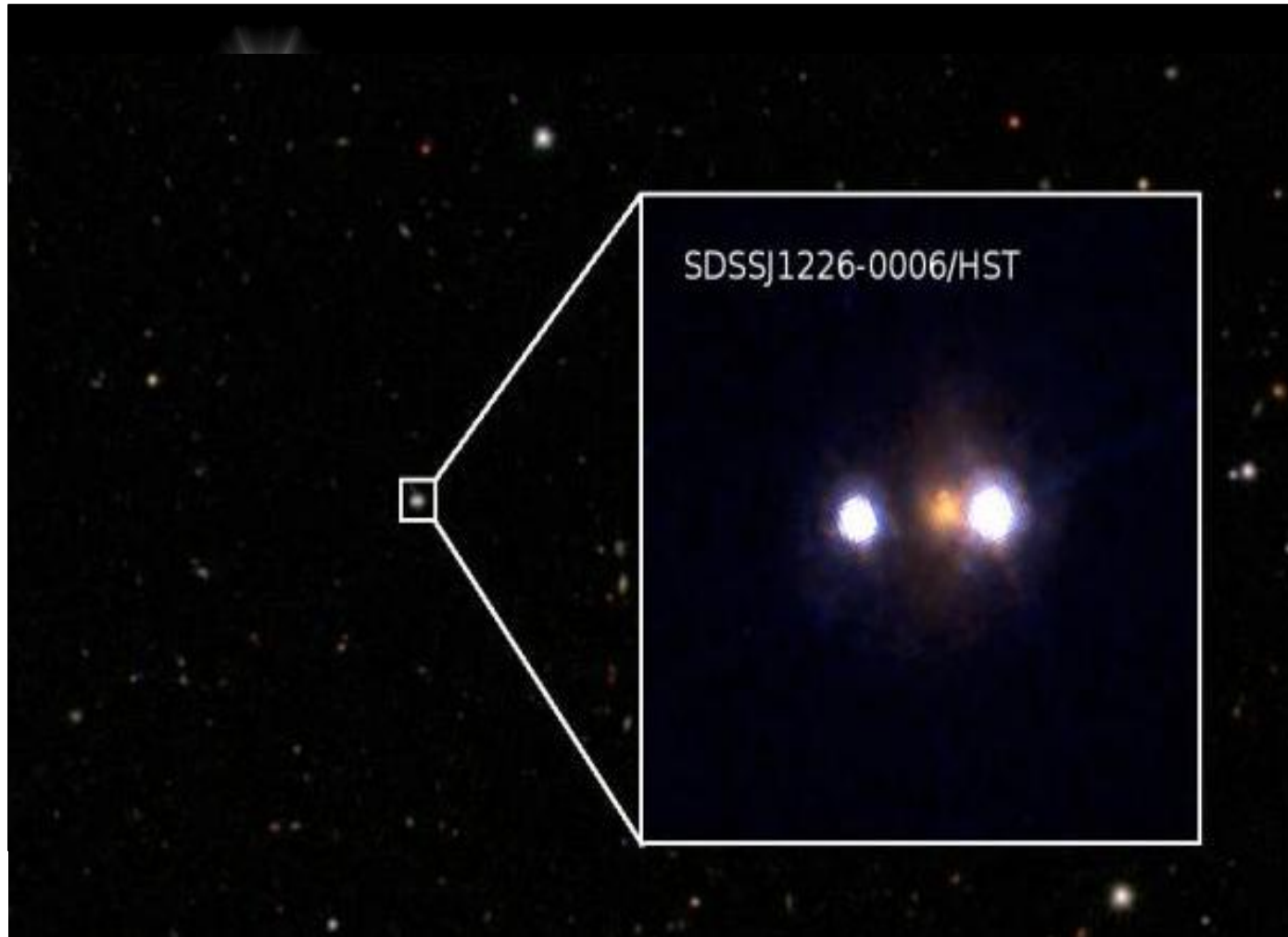
- Image of distant objects is ***distorted*** and ***magnified*** by any gravitational field it passes through.
- What does this actually look like?
- You can play with this-- with a publically available app! (for iPhone, made by KIPACker, Eli R.)
- Check out **gravlens3** on iTunes.

# Gravitational lensing: multiple images

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# Gravitational lensing: multiple images

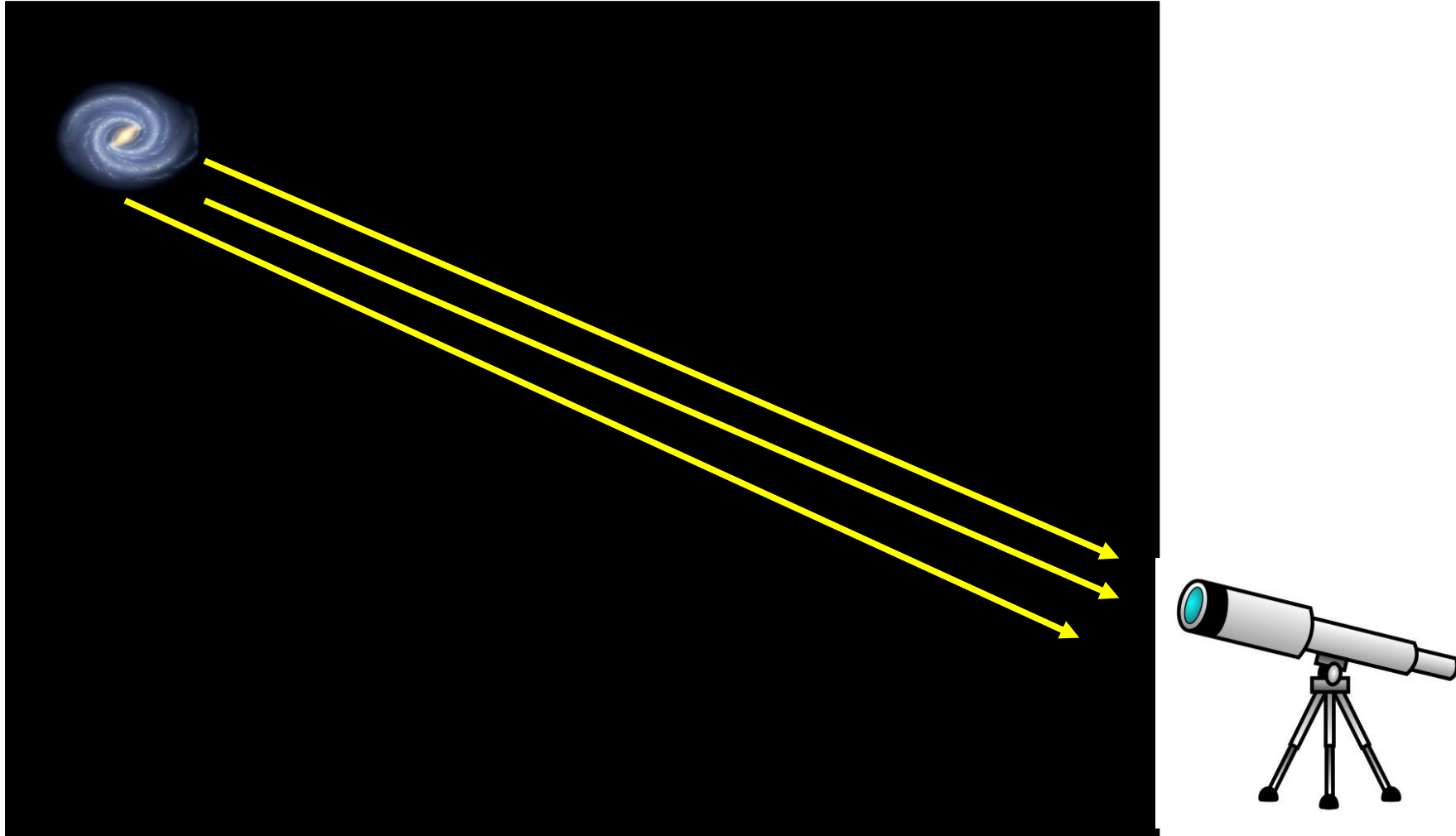


*Image Credit: Masamune Oguri, Naohisa Inada et al.*



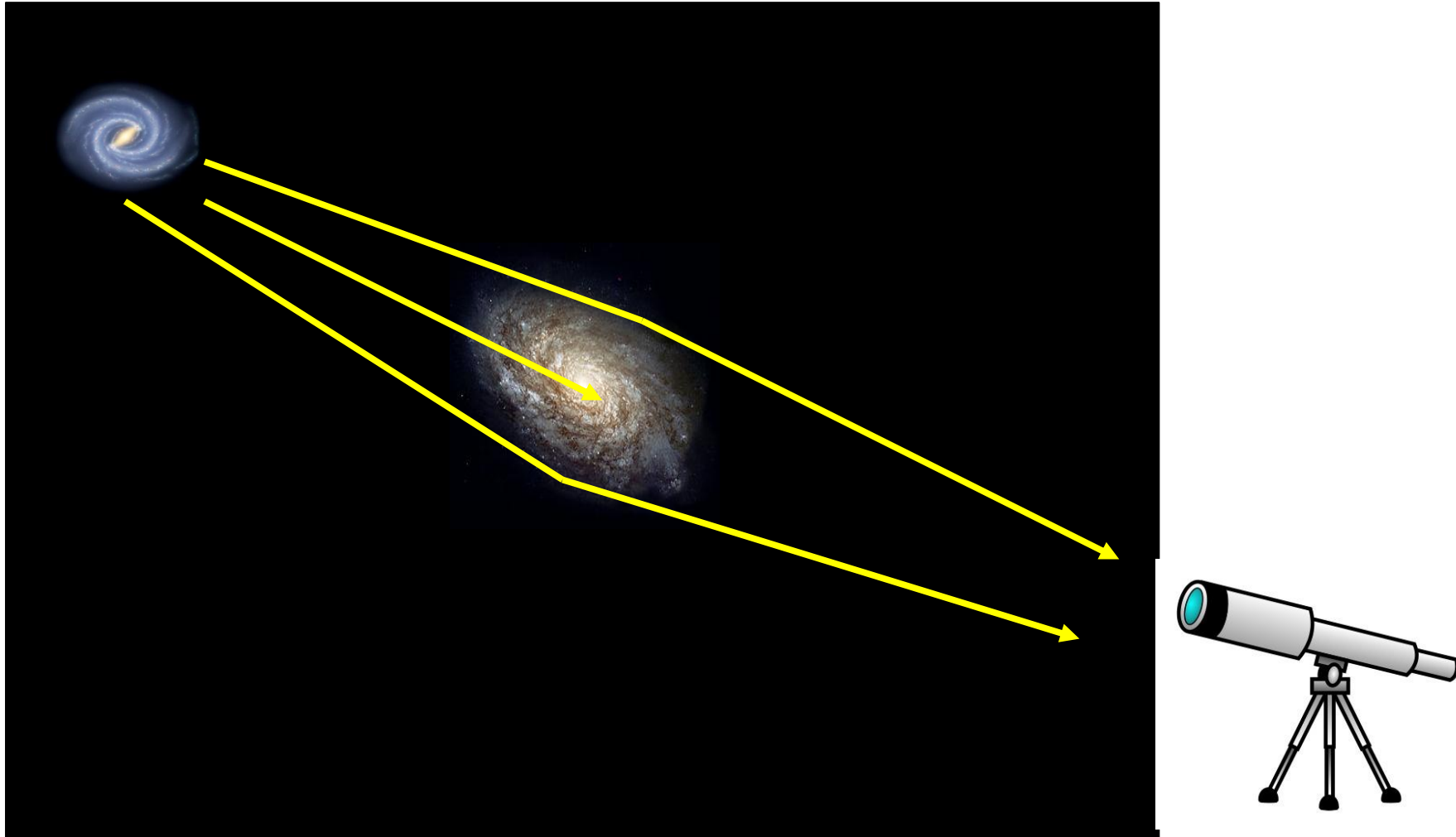
# What if you're looking at a galaxy?

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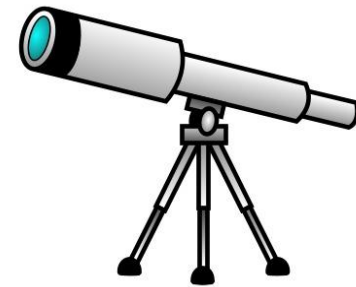
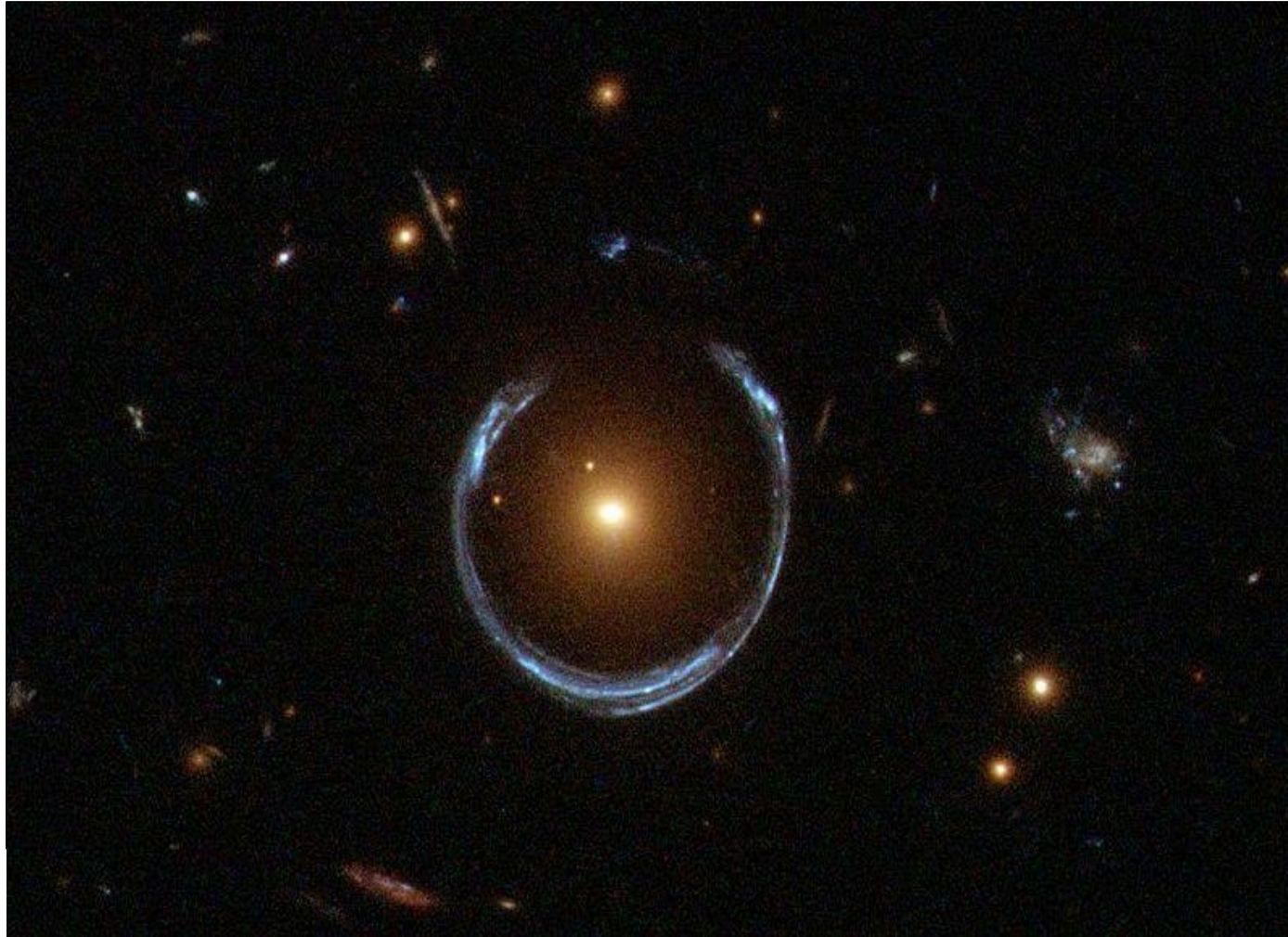


# And if there's another galaxy in the way?

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**Depending on configuration, can form an  
“Einstein Ring” (---> spectacular images!)**





# Playing with GravLens3

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Original  
image



Debbie Bard of SLAC

# Gravity *lenses*... Debbie!

(Distortion when lens is to the side front of source )

---

Lens radially  
distant from  
Source

---> Slightly  
distorted  
image!



# Distortion when lens is in front of source

---

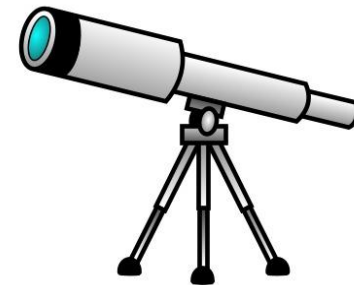
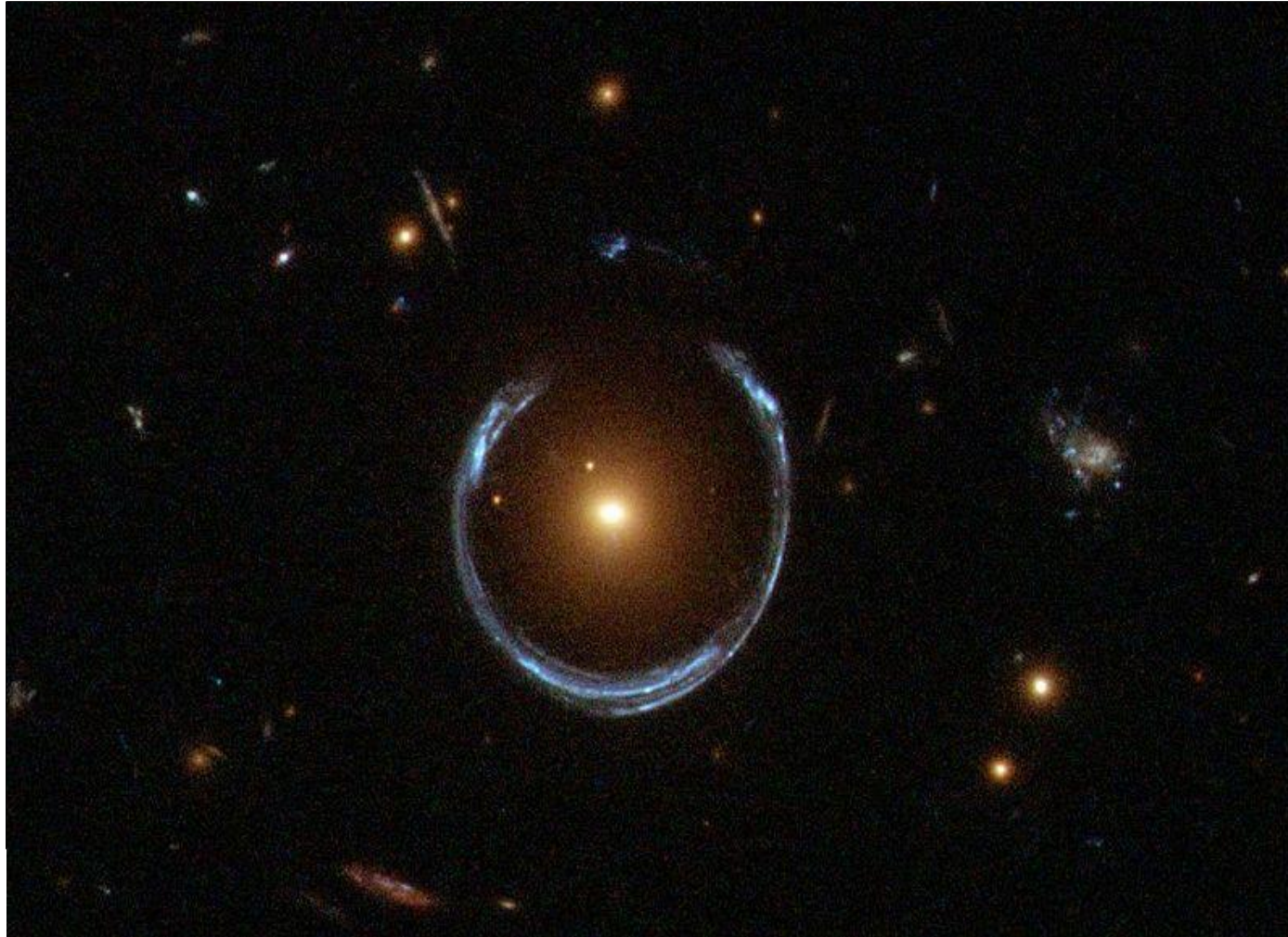
Very  
distorted  
image!





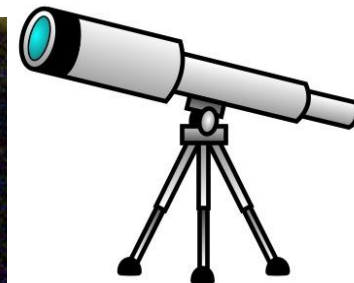
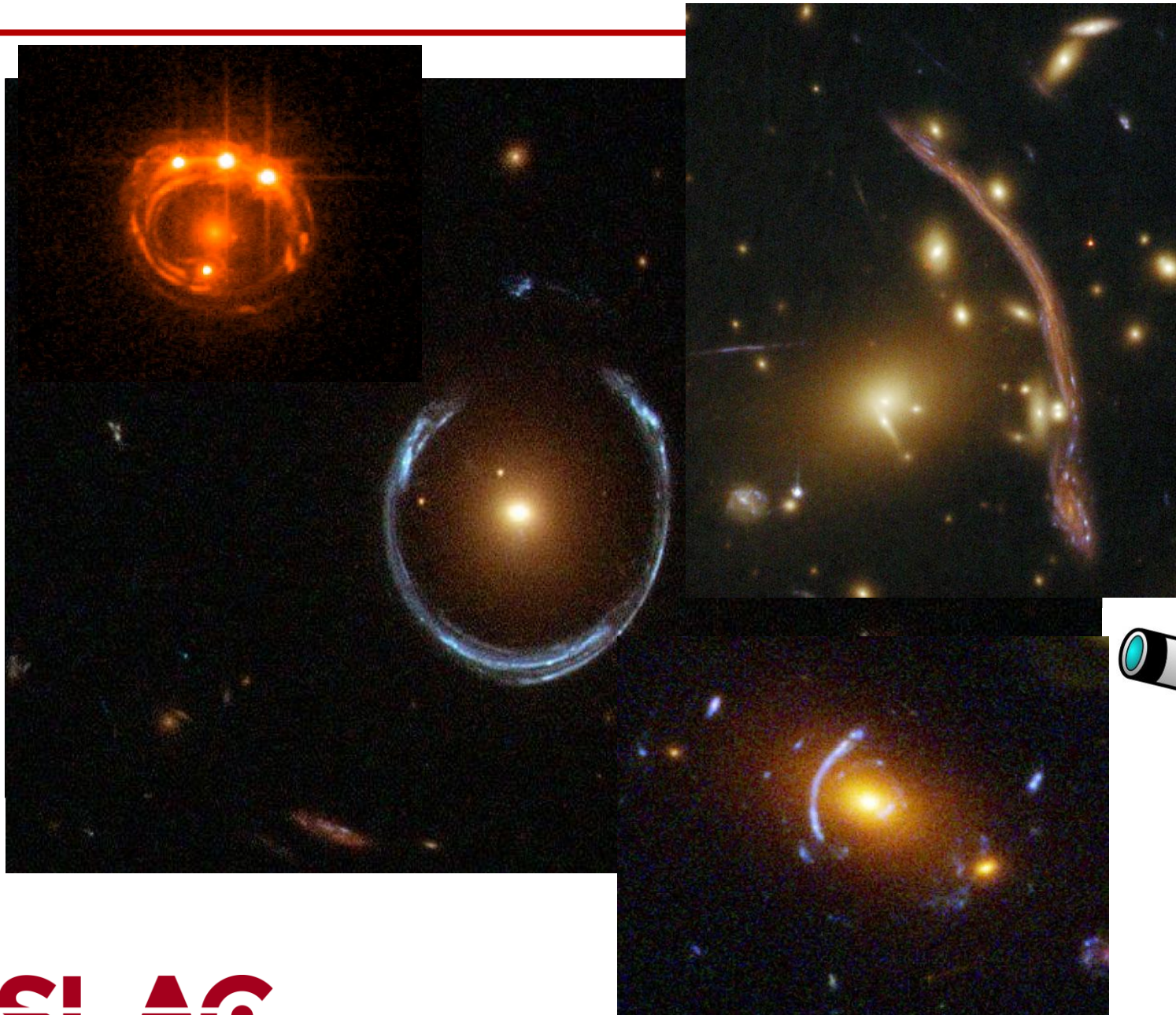
# An Einstein Ring ...and an “Einstein Eye”

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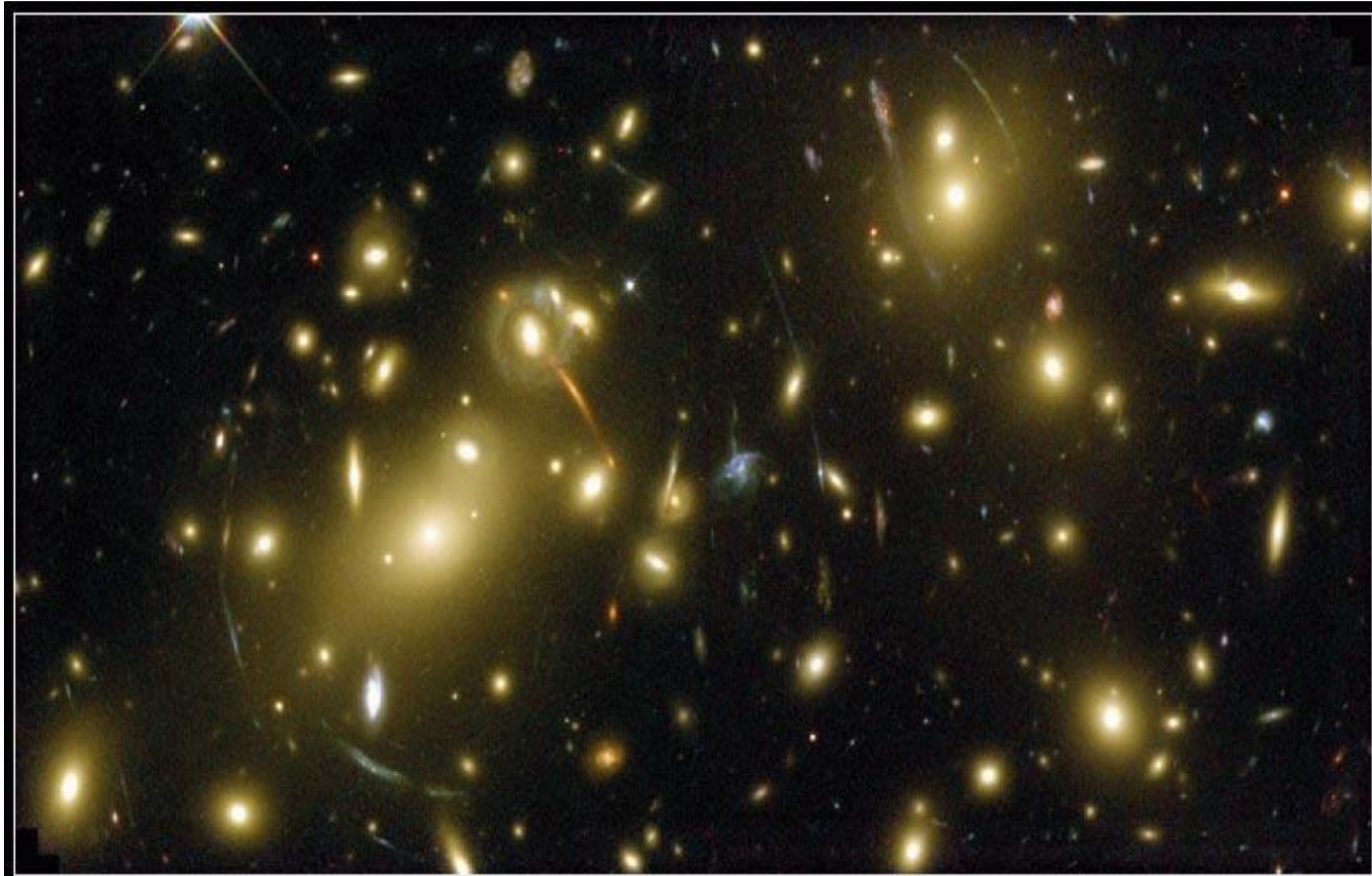


# And: Einstein Arcs



# A lot of mass (galaxy cluster) → A whole lot of lensing

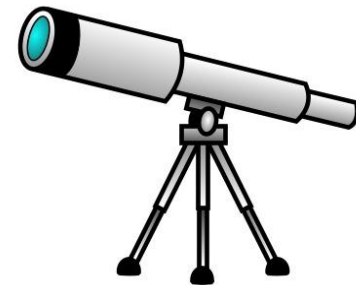
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**Galaxy Cluster Abell 2218**

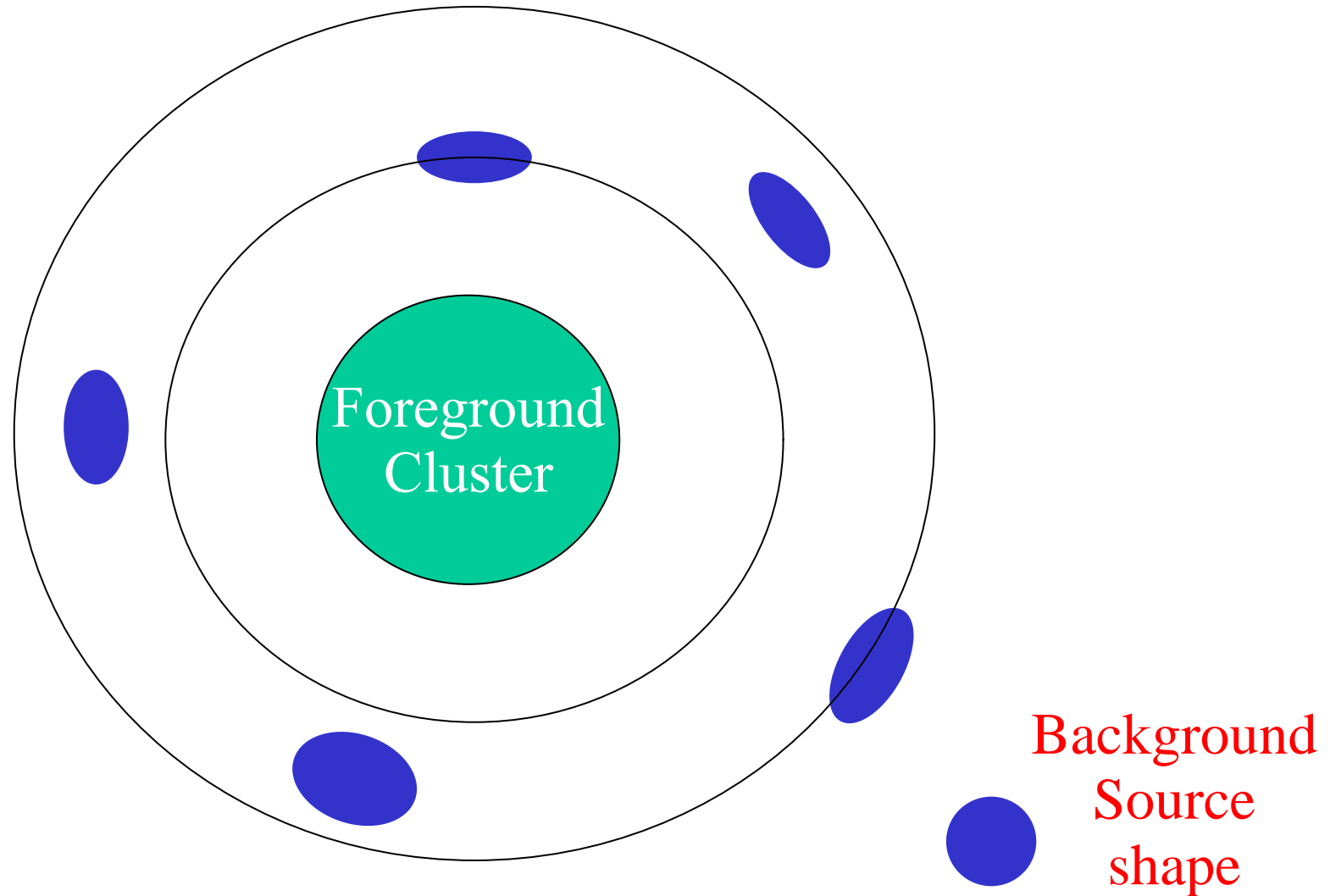
**HST • WFPC2**

NASA, A. Fruchter and the ERO Team (STScI, STECF) • STScI-PRC00-08



# Lensing Effect on Background Galaxies

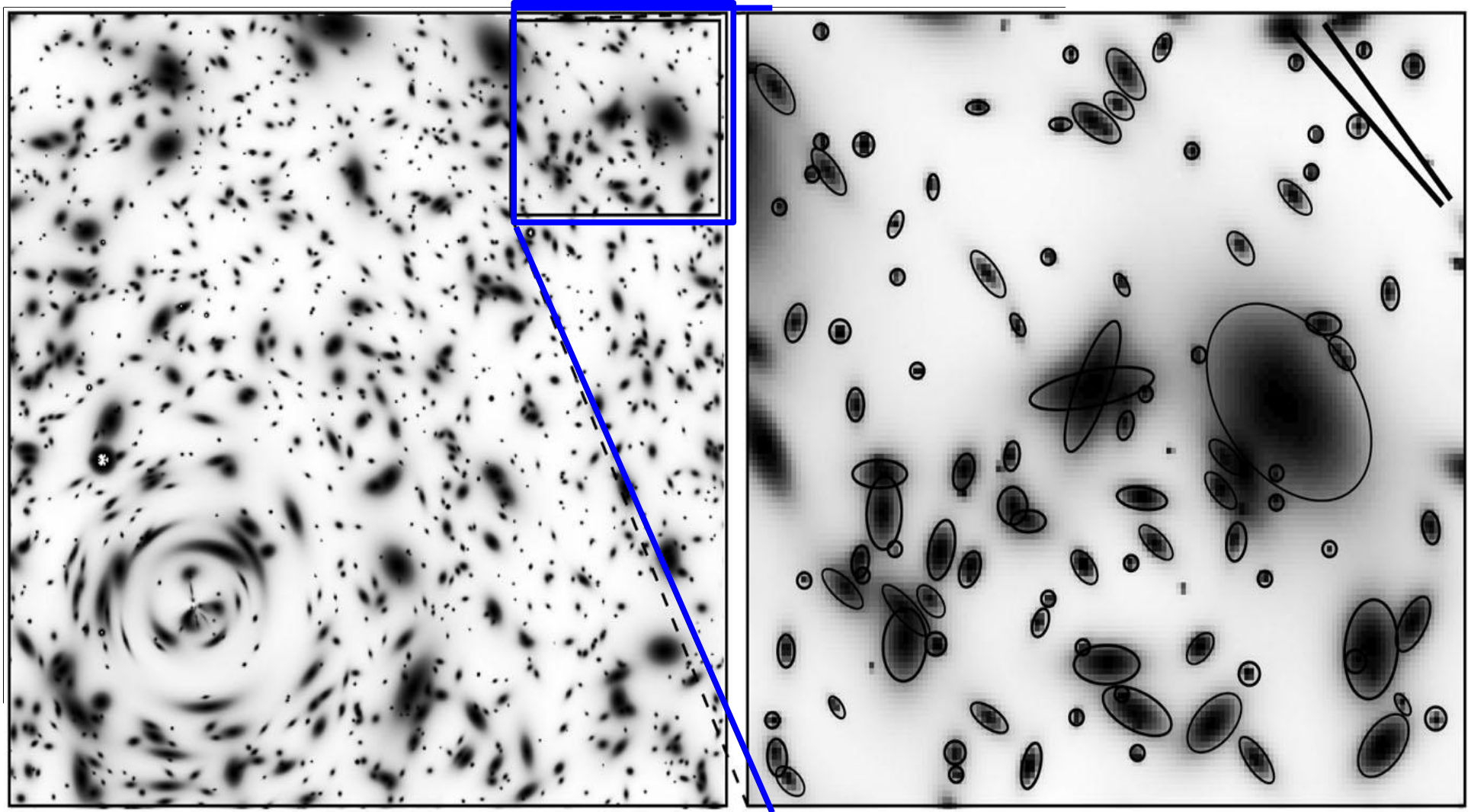
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Note: the effect has been greatly exaggerated here



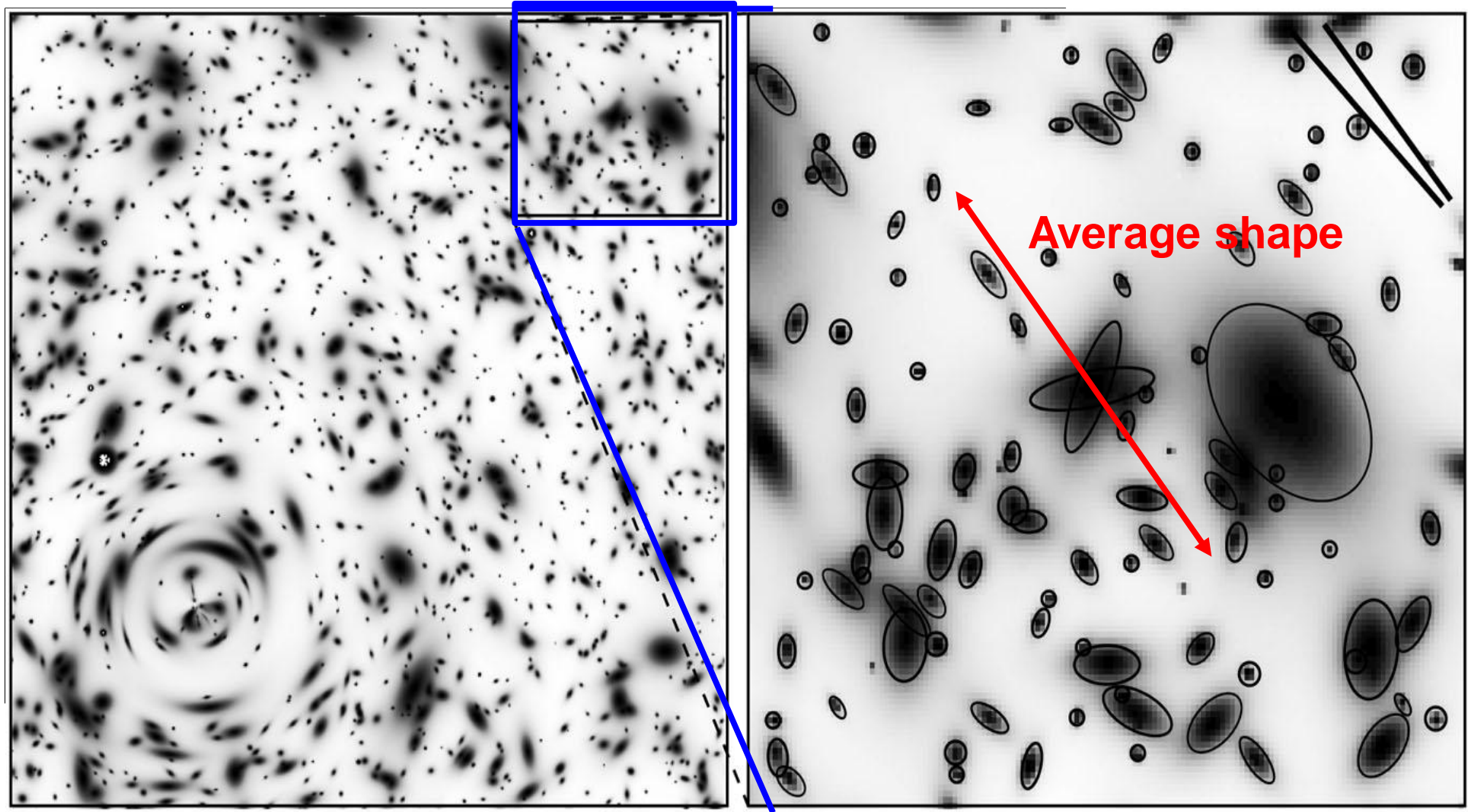
# However: 'low level' lensing is not always obvious to the eye...



*Simulation: Tyson et al.*

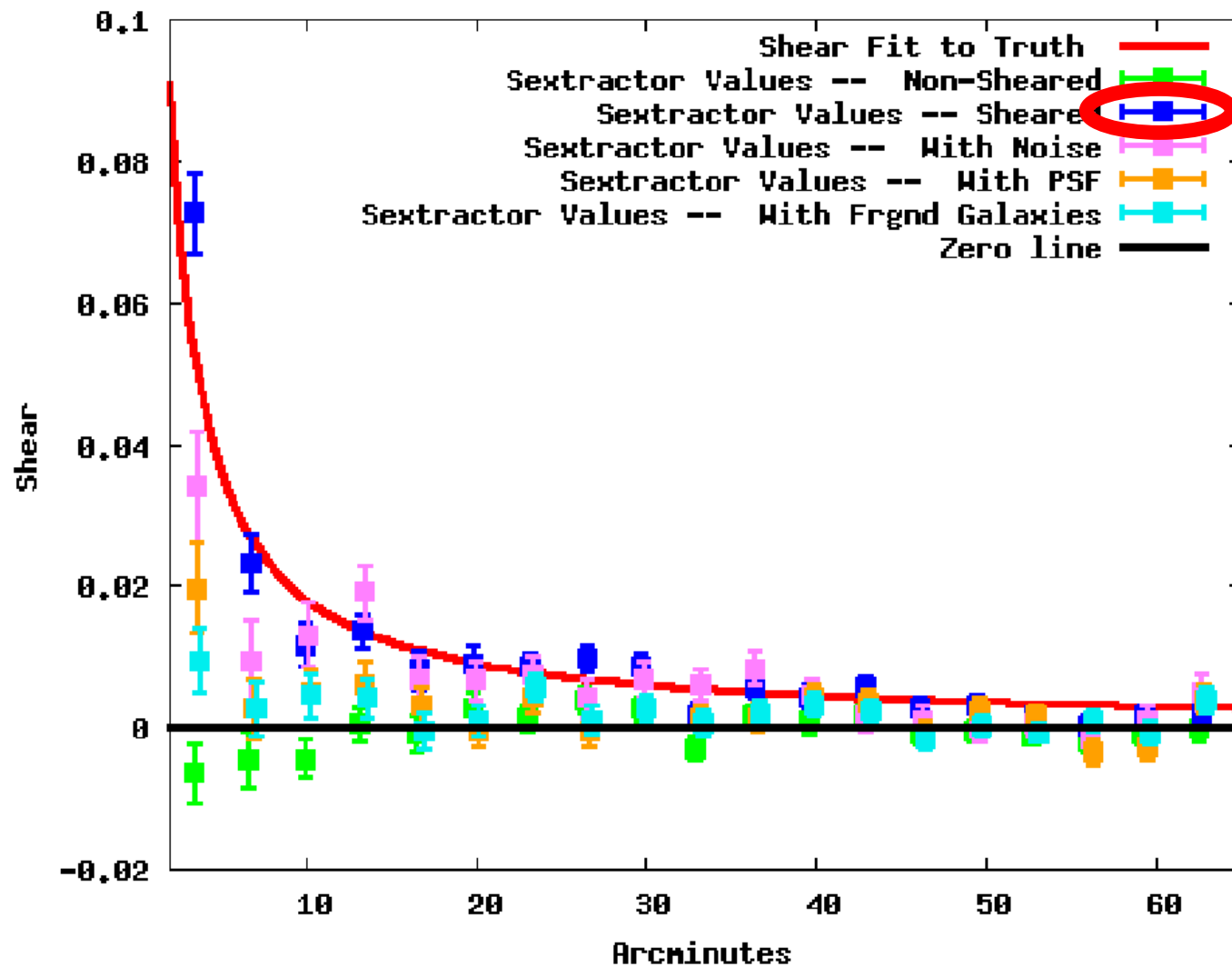


# But *when averaged*, it can be picked out



*Simulation: Tyson et al.*

# Lensing going out from lens center





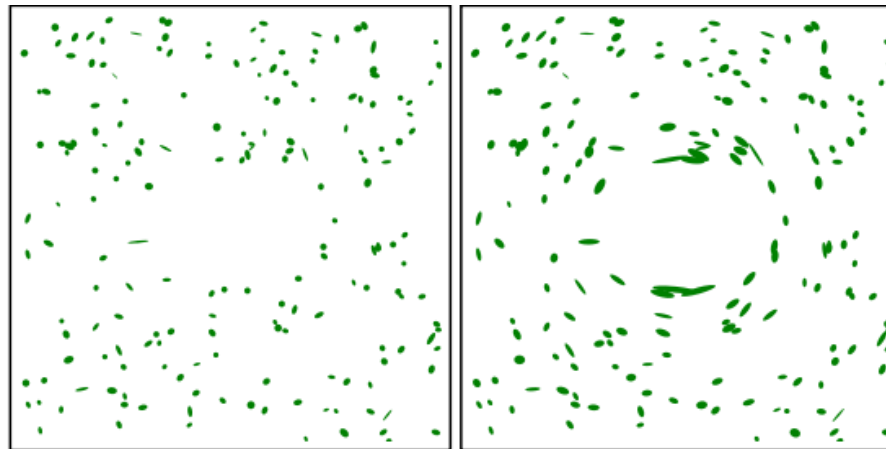
The background of the slide is a deep black space filled with a vast number of bright blue stars. The stars vary in size and brightness, with a particularly dense and bright cluster of stars located in the center of the image. The overall effect is a sense of a distant, star-filled galaxy or nebula.

# Dark matter and lensing

# Reconstructing the lens

---

- Gravity lenses light → all massive objects lens
  - Stars, galaxies, galaxy clusters... all are mass lenses!
- We can use the lensed image of distant galaxies to reconstruct the mass of the ***lens itself***



No lensing

Lensing



# Reconstructing the lens

- From a Hubble image of cluster of galaxies.
- We want to figure out how much mass there is in this cluster...



**Gravitational Lens**  
**Galaxy Cluster 0024+1654**

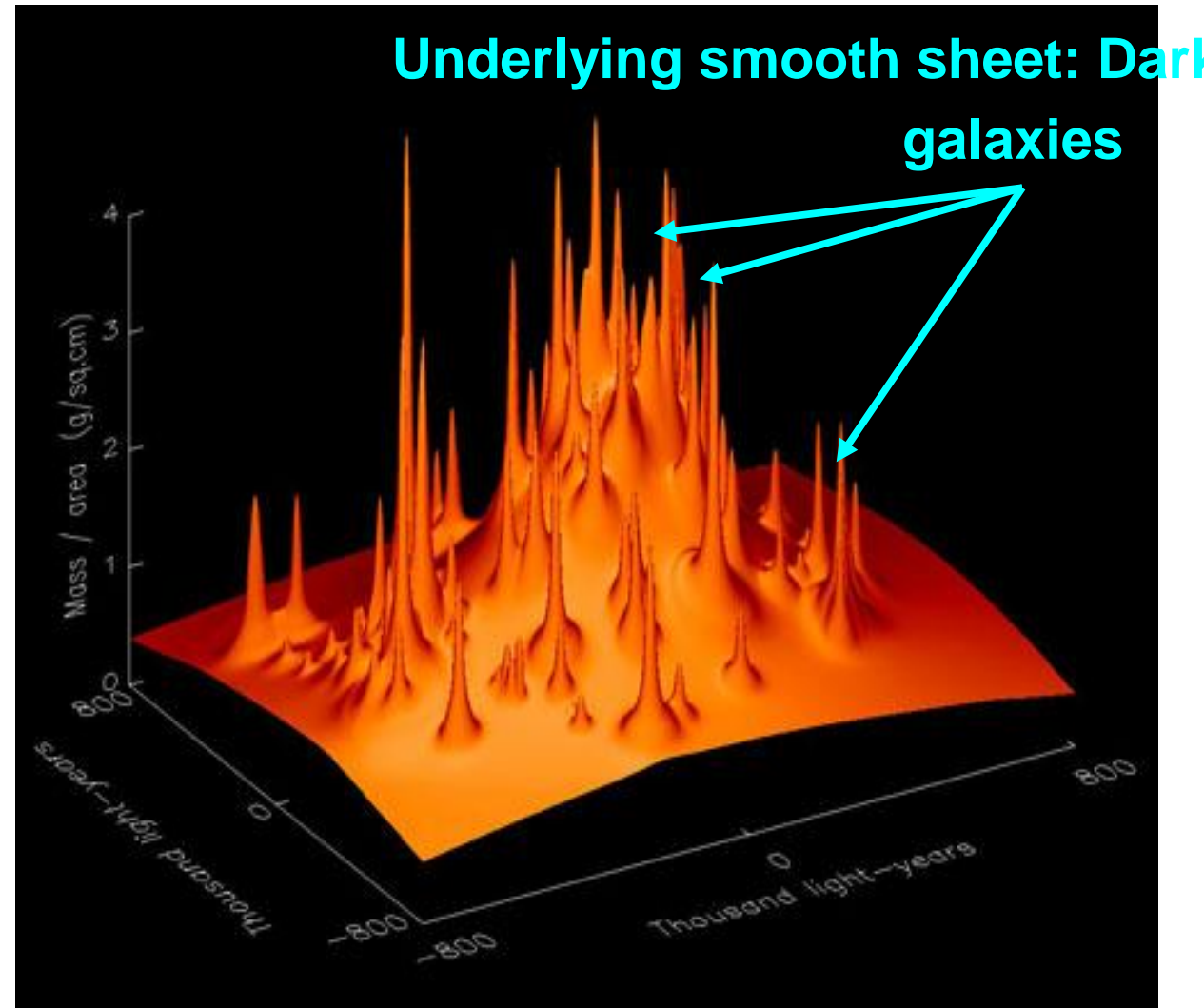
HST • WFPC2

PRC96-10 • ST ScI OPO • April 24, 1996

W.N. Colley (Princeton University), E. Turner (Princeton University),  
J.A. Tyson (AT&T Bell Labs) and NASA

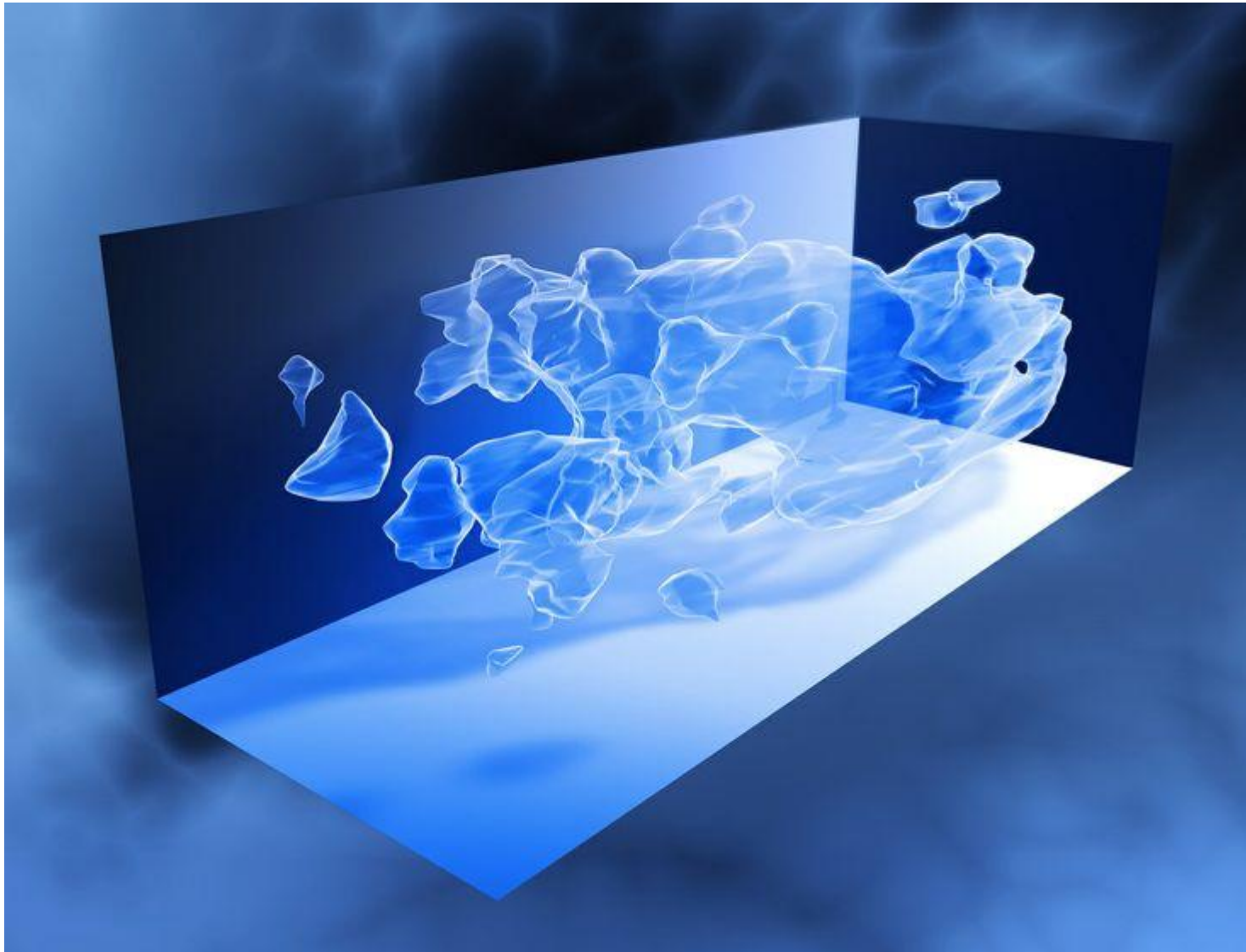
# Reconstructing the lens

- Hubble image of cluster of galaxies.
- We want to figure out how much mass is in this cluster...
- Reconstruct lensing mass, based on distortion images of background galaxies!



# Mass Map of a Part of Universe from Lensing

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A visualization of the cosmic web, showing a dense network of purple filaments and nodes of yellow and orange light against a dark background. The filaments represent the large-scale structure of the universe, while the nodes represent galaxy clusters.

# **Dark energy and lensing**

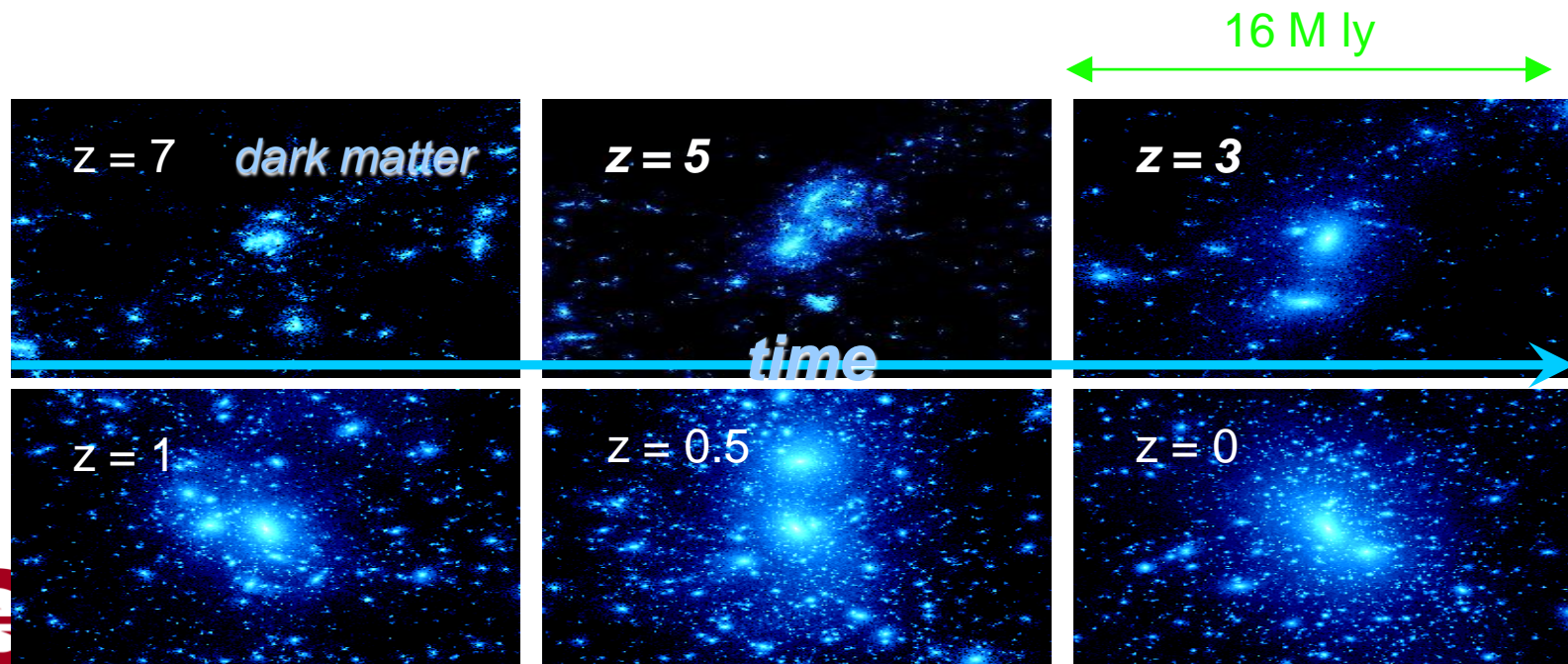


# Galaxy Cluster Growth in Cosmic Time

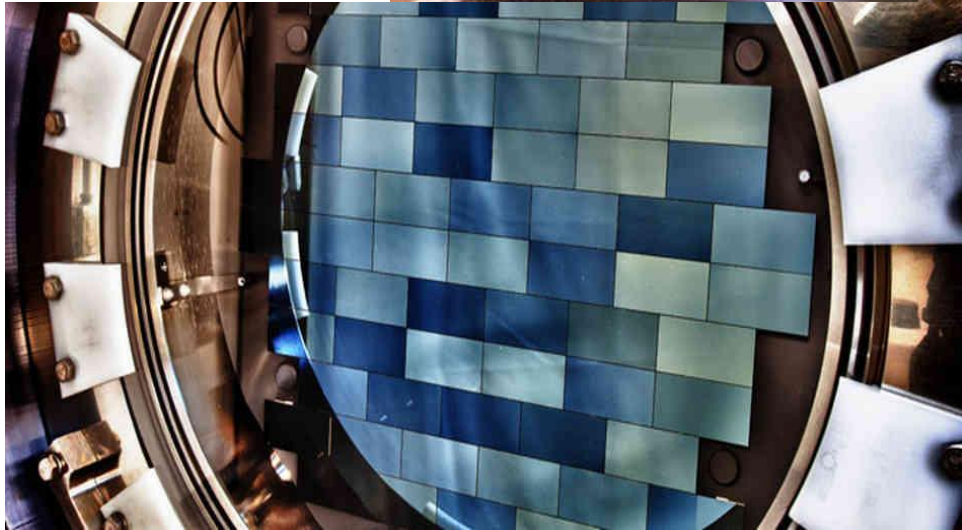
Dark Energy opposes galaxy cluster collapse

Thus we can learn about  $\Lambda$  using

mass of clusters as a function of  $z$  = redshift ( $z=0$  is now, higher  $z$  is farther back in time)

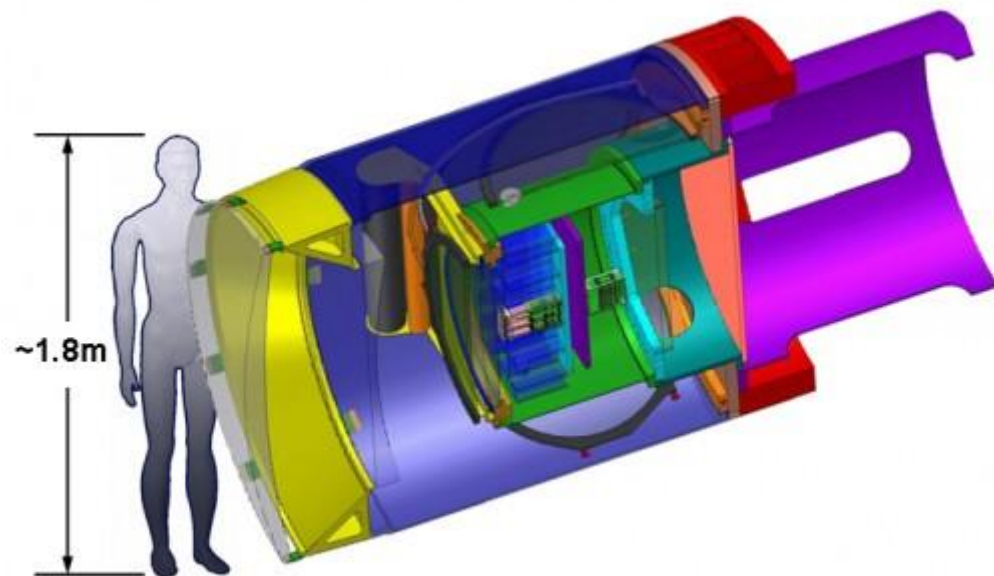
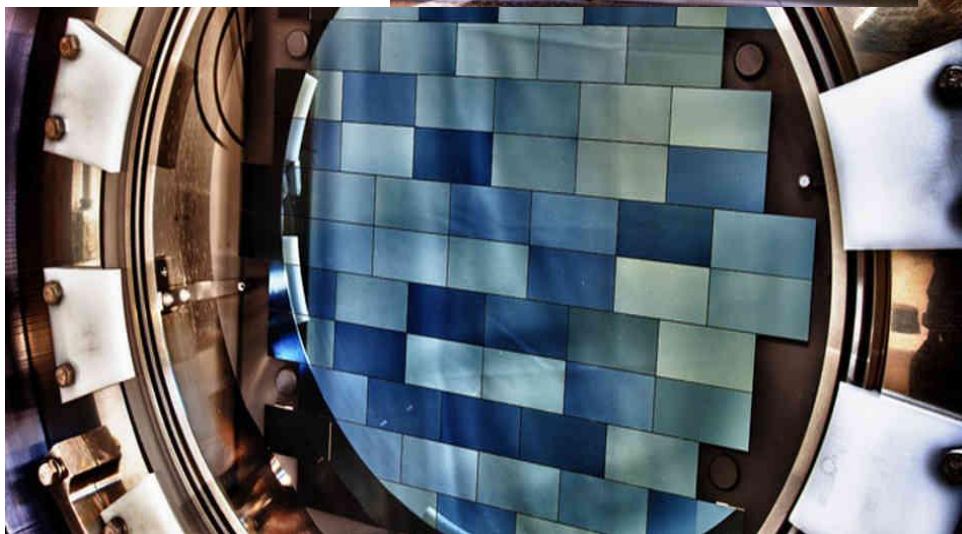


# Gravitational Lensing and SLAC





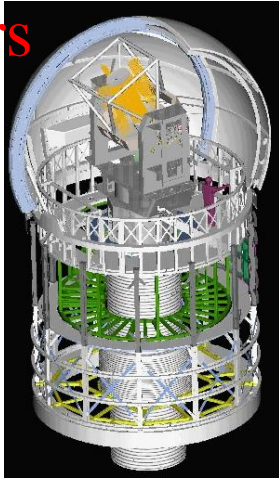
# Gravitational Lensing and



# Several Upcoming Lensing Surveys

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Panstarrs



Dark Energy Survey

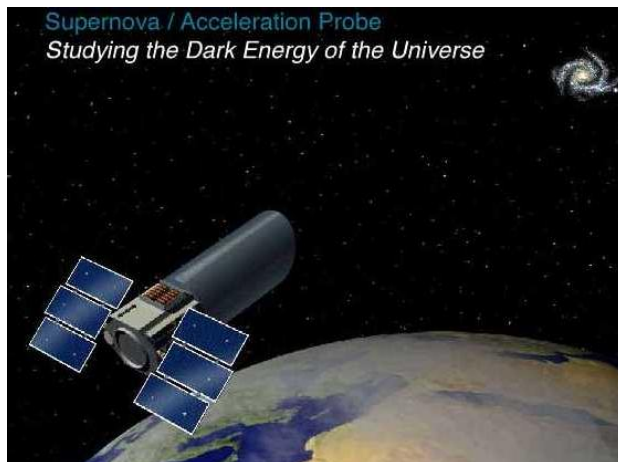


DARK ENERGY  
Survey

LBT-OWLS

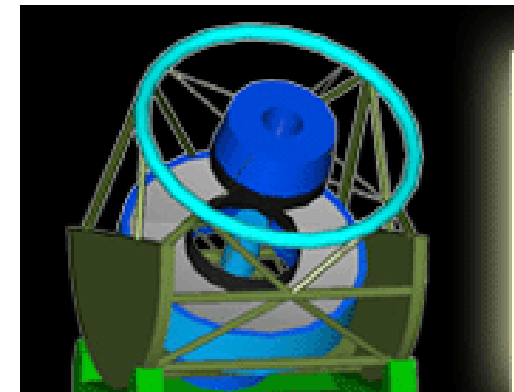


SNAP



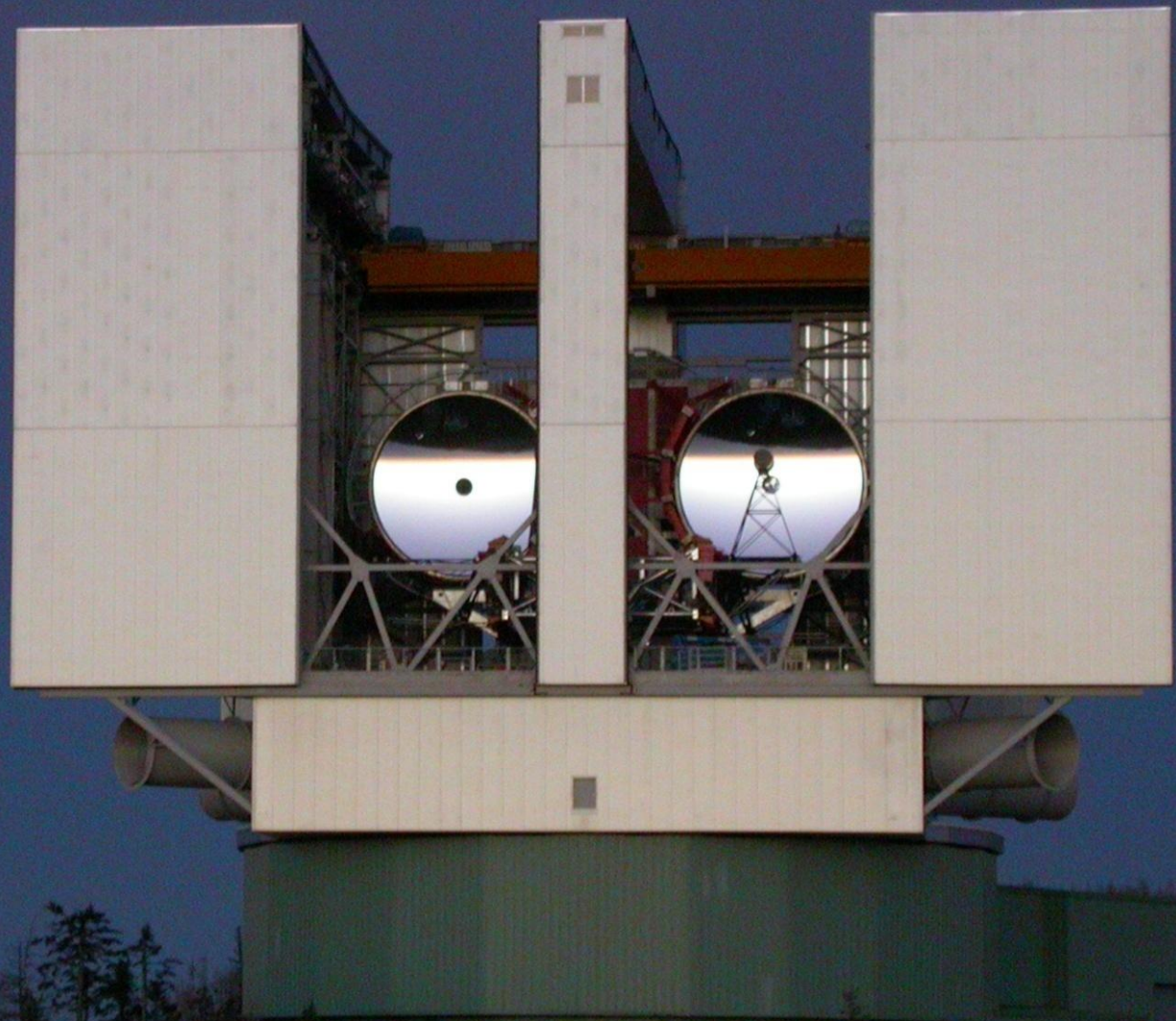
NATIONAL ACCELERATOR LABORATORY

LSST





# The Large Binocular Telescope



World's largest optical telescope (two primary mirrors, each 27 feet in diameter)



# Summary (1/2)

- **Lensing is an amazing probe of matter in the universe.**
- **Galaxy clusters give some of the strongest evidence for dark matter.**
- **Lensing can constrain theories of dark energy.**




# Summary (2/2)

- **Dark Energy Survey and Large Synoptic Survey Telescope allow us to probe deeper than ever before.**
- **And, on the interactive side...**



# Some things to play with...

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-  **GravLens3** (Eli Rykoff)
  - Free app on the iTunes store
- **spacewarps.org** (Phil Marshall)
  - Crowd-sourcing strong lens identification

# Finally: How a Cosmic Perspective Can be valuable to us in our Lives

*To my mind, there is perhaps no better demonstration of the folly of human conceits than this distant image of our tiny world. To me, it underscores our responsibility to deal more kindly and compassionately with one another and to preserve and cherish that pale blue dot, the only home*

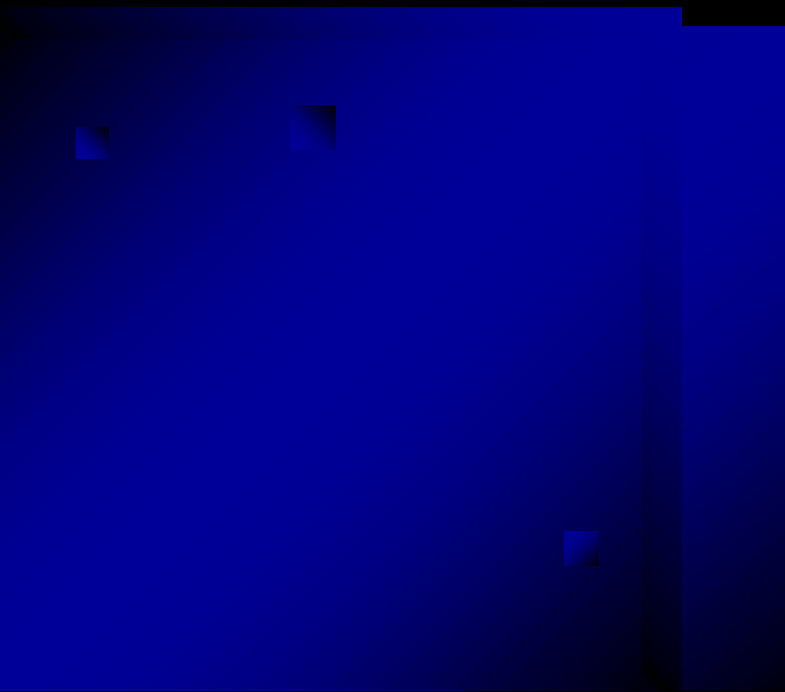
*we've ever known.*

Carl Sagan, May 1996

See more at: <http://www.skyimagelab.com/pale-blue-dot.html#sthash.zkPOhAkx.dpuf>

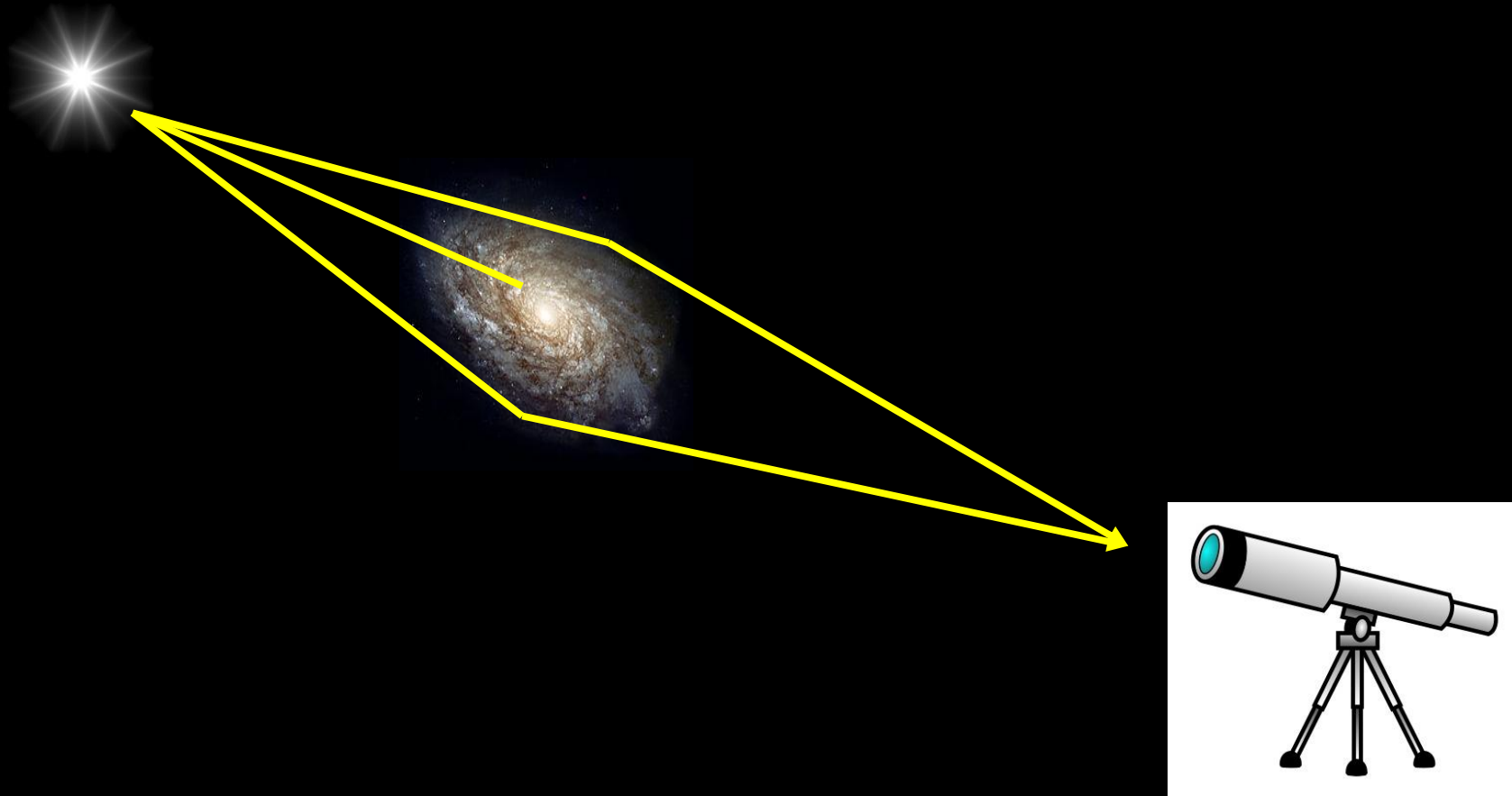
*I maintain that the cosmic religious feeling is the strongest and noblest motive for scientific research – A. Einstein, NYT Magazine, 1930*

# Backup Slides



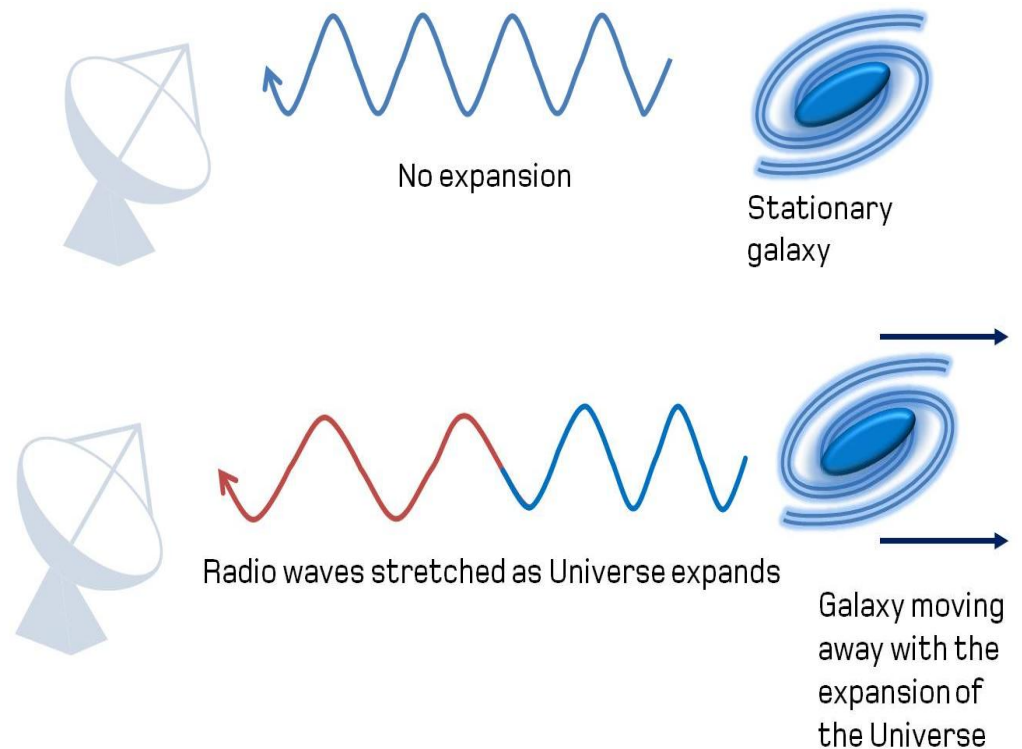


## Gravitational lensing: magnification



# The Universe is expanding

- Redshift: like Doppler effect, but with light.
- Light from objects moving away from us appears more **red**.
- Universe is expanding → more distant objects appear more **red** to us.



# A Very Quick History of Particle Physics

1869 Periodic table finalized – all elements fit

1897  $e^-$  = electron (UK)

1910's proton (UK etc.)

1932 neutron (UK)

By 1932: tidy picture of three basic particles explained everything in the Universe

1932  $e^+$  = positron (cosmic rays -- CA)

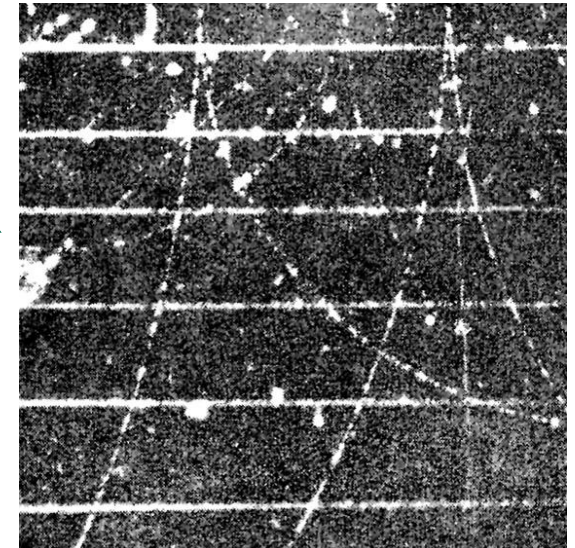
1936  $\mu^-$  – (c. rays – CO and Panama )

1947  $\pi$  (c. rays – France and Bolivia )

1947 Kaon (c. rays – Mt. Wilson, CA)

1947 Lambda (c. rays – UK)

Then: many, many more...





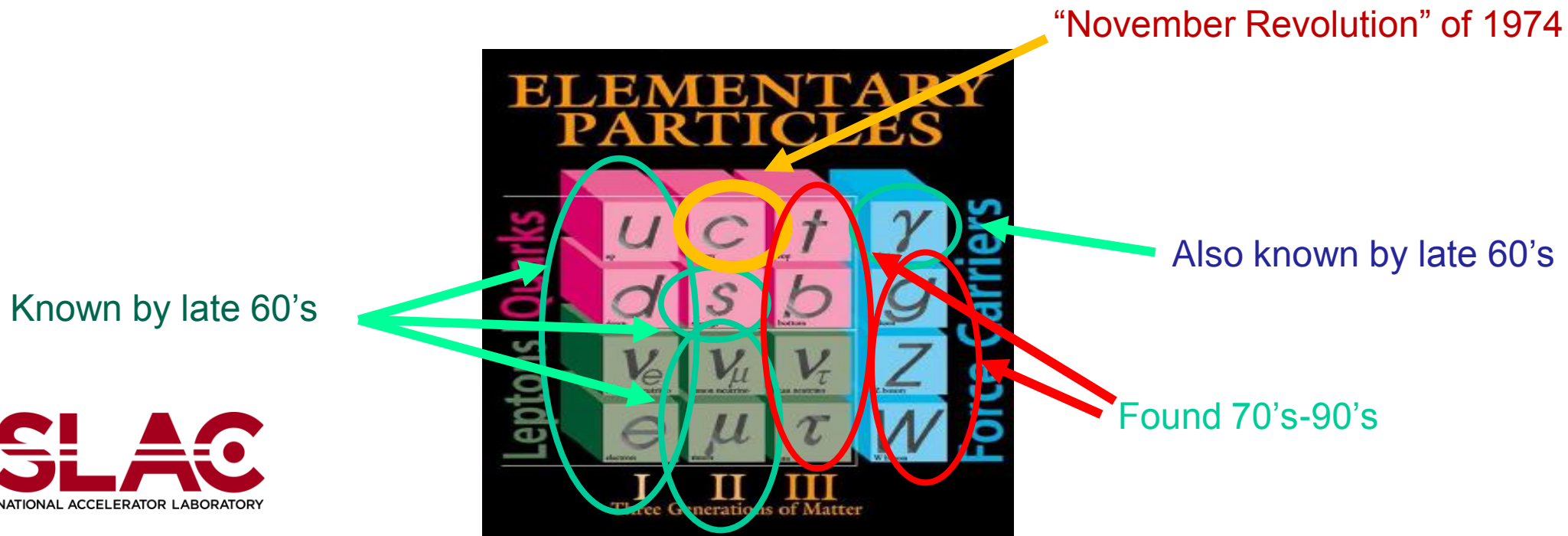
# Finally: the “congealing” of the SM (The Standard Model)

Late 60's: Electroweak theory combines electromagnetism and weak nuclear force

Then: quarks + gluons = nuclear physics ironed out

1970's-1990's: all other predicted SM particles tracked down (heavier quarks and leptons, more force carriers)

All new matter particles fit as an extra generation into the already existing theory



# Putting together the building blocks

Anything made of quarks is a **Hadron**

## **Baryons** $qqq$ and Antibaryons $\bar{q}\bar{q}\bar{q}$

Baryons are fermionic hadrons.

These are a few of the many types of baryons.

Symbol	Name	Quark content	Electric charge	Mass $\text{GeV}/c^2$	Spin
<b>p</b>	proton	<b>uud</b>	1	0.938	1/2
$\bar{p}$	antiproton	$\bar{u}\bar{u}\bar{d}$	-1	0.938	1/2
<b>n</b>	neutron	<b>udd</b>	0	0.940	1/2
$\Lambda$	lambda	<b>uds</b>	0	1.116	1/2
$\Omega^-$	omega	<b>sss</b>	-1	1.672	3/2

## **Mesons** $q\bar{q}$

Mesons are bosonic hadrons

These are a few of the many types of mesons.

Symbol	Name	Quark content	Electric charge	Mass $\text{GeV}/c^2$	Spin
$\pi^+$	pion	<b><math>u\bar{d}</math></b>	+1	0.140	0
<b><math>K^-</math></b>	kaon	<b><math>s\bar{u}</math></b>	-1	0.494	0
$\rho^+$	rho	<b><math>u\bar{d}</math></b>	+1	0.776	1
<b><math>B^0</math></b>	B-zero	<b><math>d\bar{b}</math></b>	0	5.279	0
$\eta_c$	eta-c	<b><math>c\bar{c}</math></b>	0	2.980	0

How to make up hadrons from quarks

Now we see it's back to being simple! (relatively.. ☺ )